

# How far can we see back in time in high-energy collisions using charm quarks?

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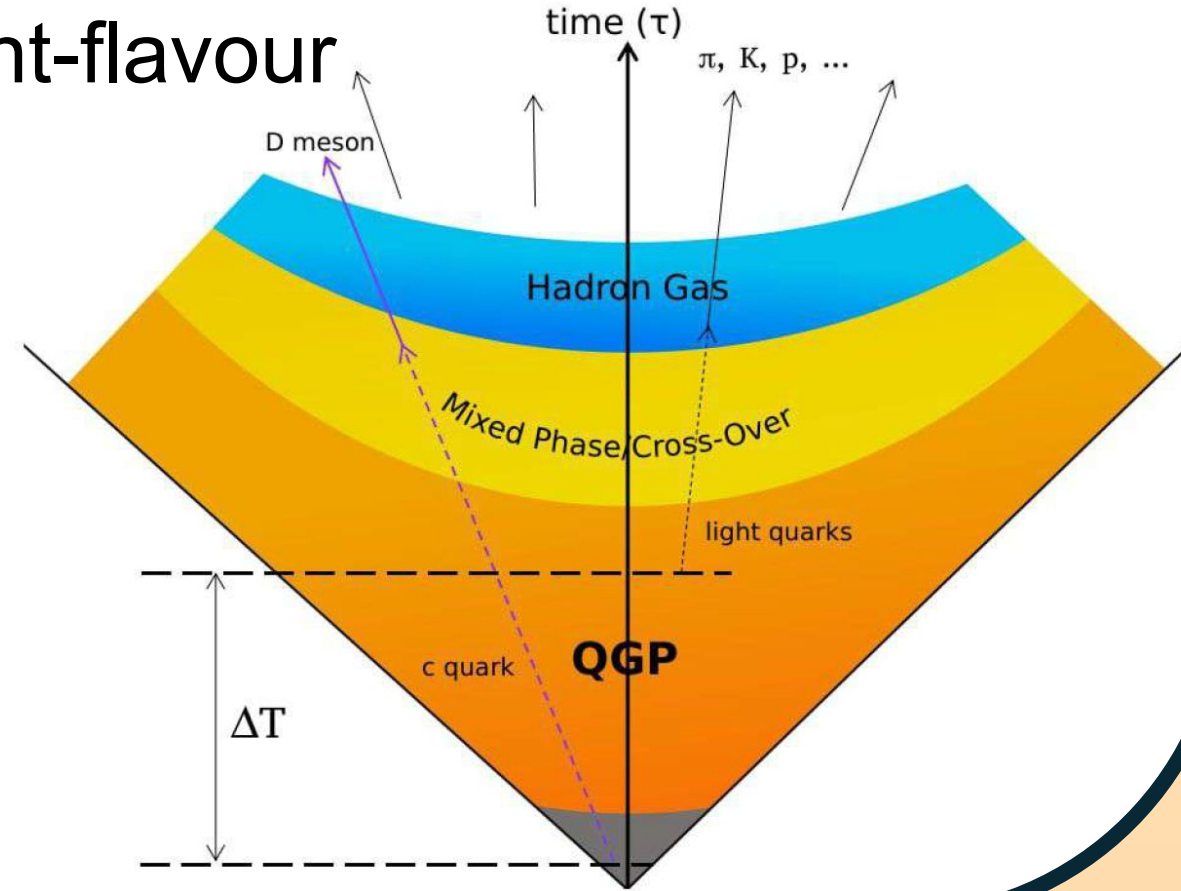
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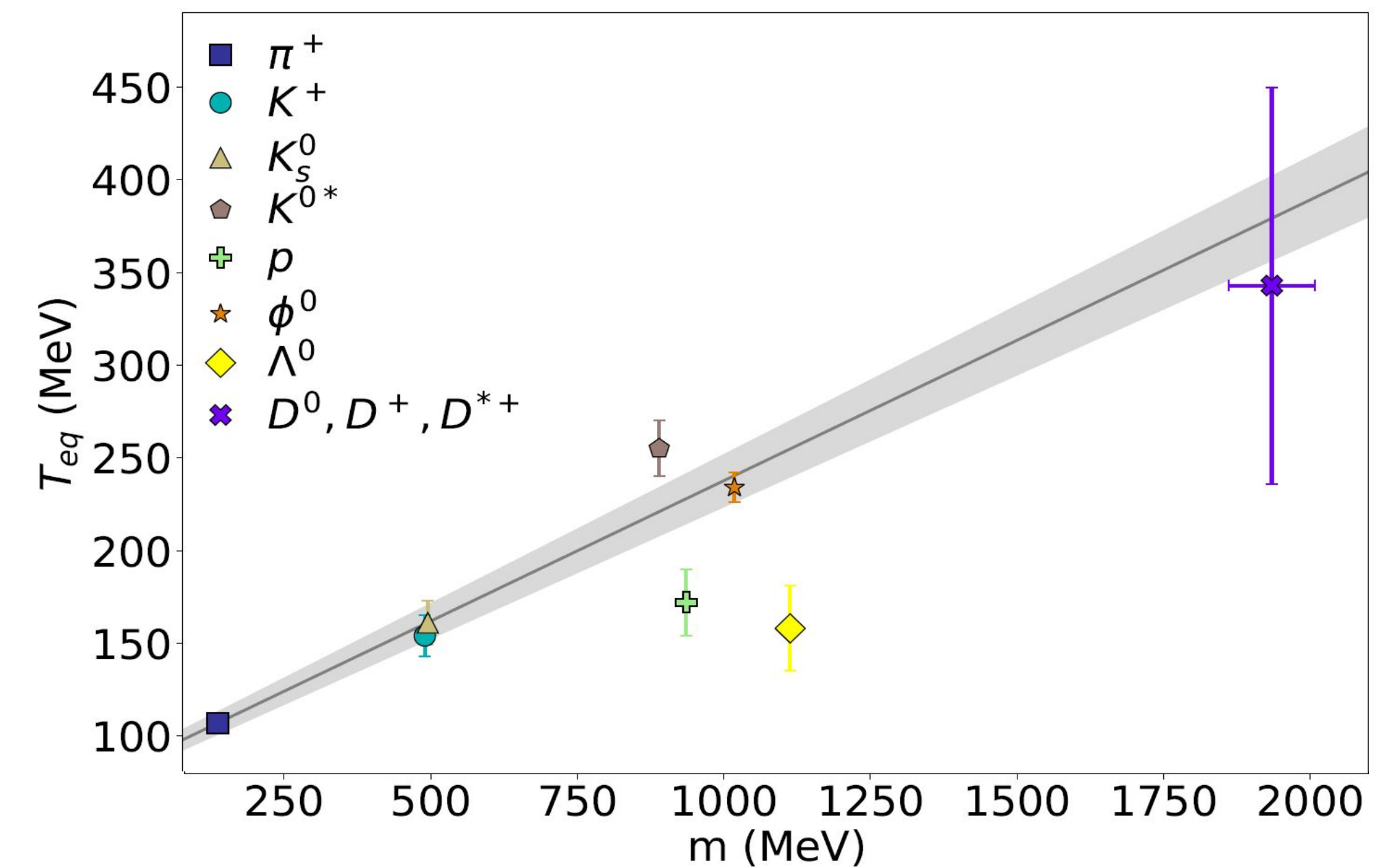
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## Motivation

- In high-energy collisions, light-flavour and strange hadrons mostly carry information about the final state
- Heavy-flavour  $c$  and  $b$  quarks are produced in the initial stages of a collision and experience the whole evolution of the system
- Measurement of heavy-flavour hadron production allows for studying the earlier stages of a collision
- The non-extensive Tsallis – Pareto statistical framework[1] has been shown to describe well the spectra of light-flavour hadrons[2]
- In this work we evaluate applicability of non-extensive thermodynamical principles on heavy-flavour production
- We use D mesons to investigate the thermodynamical properties of earlier stages of the system



## Common Tsallis temperature



- The common Tsallis temperature increases with the mass of **mesons**
- The common Tsallis temperature of **baryons** is lower compared to the mesons with similar masses

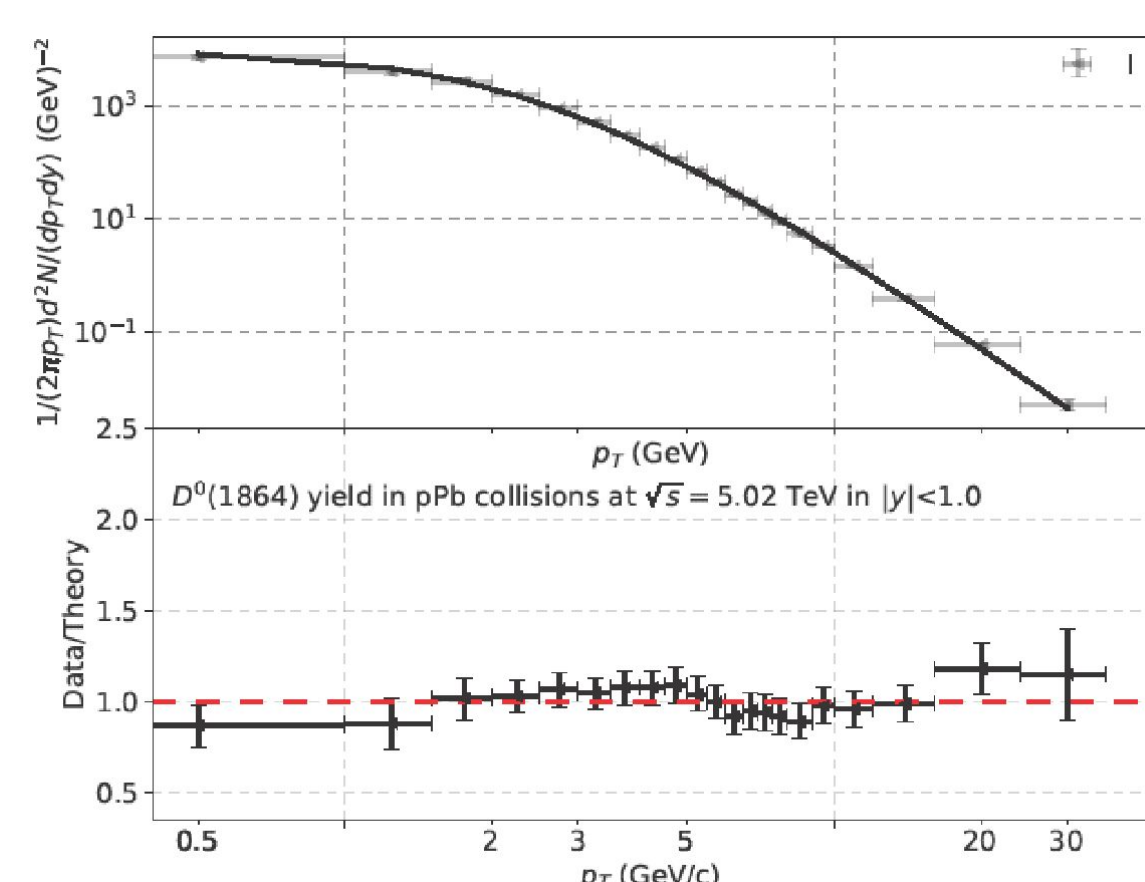
## Method

### Low- $p_T$ part of spectrum:

- Boltzmann – Gibbs distribution
- characterized by the kinetic freeze-out temperature

### High- $p_T$ part of spectrum:

- power-law distribution
- perturbative QCD hadron production



The Tsallis – Pareto distribution, motivated by non-extensive thermodynamics, provides a unified description of the full spectrum:

$$\frac{d^2N}{2\pi p_T dp_T dy} \Big|_{y \approx 0} \equiv A m_T f^q = A m_T \left[ 1 + \frac{q-1}{T} (m_T - \mu) \right]^{-\frac{q}{q-1}}$$

## Spectra used in this analysis

### STAR experiment:

- $D^0$  in Au–Au collisions at  $\sqrt{s_{NN}} = 200$  GeV

### ALICE experiment:

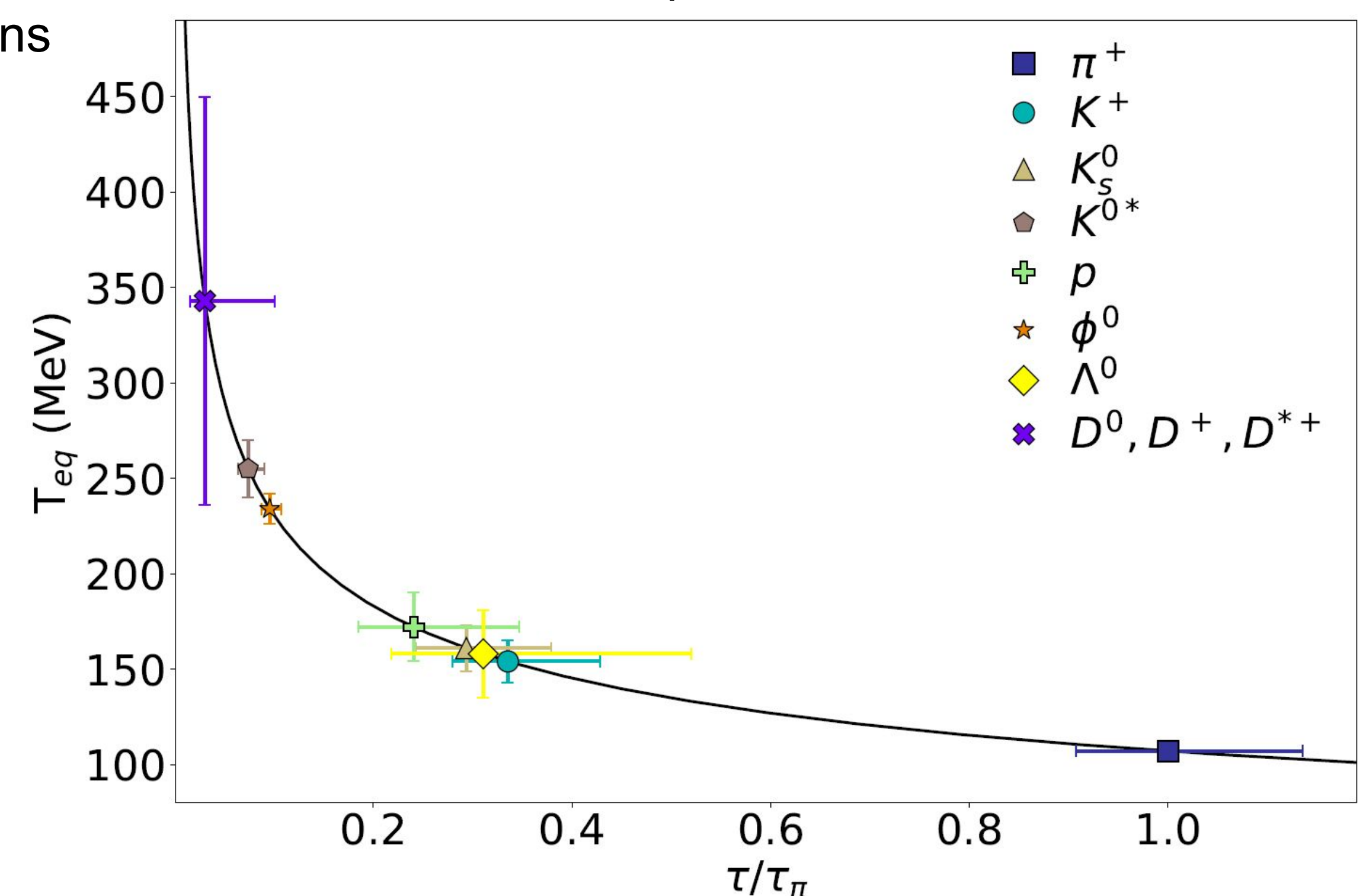
- $D^0$  in Pb–Pb collisions at  $\sqrt{s_{NN}} = 2.76$  TeV
- $D^0$ ,  $D^+$  and  $D^{*+}$  in p–Pb collisions at  $\sqrt{s_{NN}} = 5.02$  TeV
- $D^0$ ,  $D^+$  and  $D^{*+}$  in pp collisions at  $\sqrt{s} = 5.02$  TeV
- $D^0$ ,  $D^+$  and  $D^{*+}$  in pp collisions at  $\sqrt{s} = 7$  TeV

Light-flavour and strange hadron spectra are from references in [2]

## How far can we see back in time?

A simple Bjorken model can be utilised as the expansion mechanism of the ideal, ultra-relativistic matter. It yields the relation between temperature and proper time (cooling curve) [3]:  $\tau = \tau_0 \left( \frac{T_0}{T} \right)^3$

- The spectrum of pions forms much later compared to other hadrons
- Baryons are formed later compared to mesons with similar masses
- The formation time of D-meson spectra is  $\sim 30$  times shorter than for pions

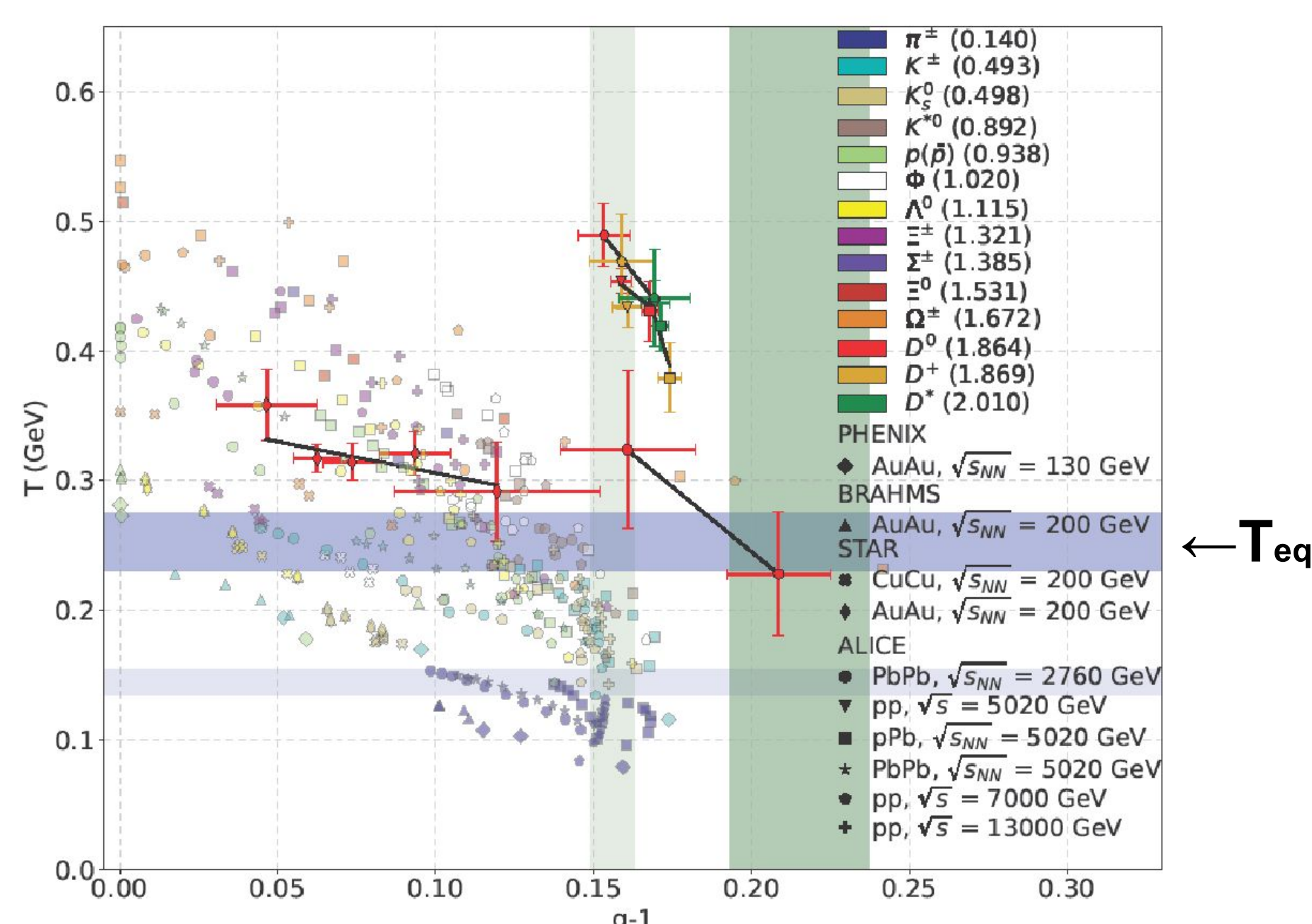


## Summary

- Transverse momentum distributions of heavy-flavour D mesons are well described by the Tsallis – Pareto distribution motivated by non-extensive thermodynamics
- The Tsallis parameters of the fits to D-meson data exhibit a scaling behaviour with charged particle multiplicity and with the collision energy
- The  $T_{eq}$  parameter for D mesons is higher compared to the light flavours. Coming from a much hotter state of the system, D mesons preserve this information, unlike the light-flavour hadrons
- Production of D mesons corresponds to a significantly earlier proper time than light-flavour hadrons. The formation time of meson spectra is also mass-dependent

## Tsallis-thermometer

The  $T$  and  $q$  parameters, extracted from the D-meson fits, are presented in the „Tsallis-thermometer”,  $T - (q-1)$  diagram



### Observations:

- mass hierarchy ( $T$  increases with particle mass and multiplicity)
- grouping based on the center-of-mass energy and collision system
- in small systems  $c$  quarks come directly from the early stages of the collisions, corresponding to high  $T$  values
- grouping of all the hadrons at **small multiplicities** around specific common  $T_{eq}$  and  $q_{eq}$  values

## References

- [1] Eur.Phys.J.A 40 (2009) 257-266
- [2] J. Phys. G, 47(10):105002, 2020
- [3] Phys. Rev. C, 97(6):064903, 2018

## Acknowledgements

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## Full papers



arXiv:2401.14282



arXiv:2409.01085