

Can rotation solve the Hubble puzzle?

Particles and Plasmas in Strong Fields



G.G. Barnaföldi, I. Szapudi, I.F. Barna, B. Szigeti

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Refs.: *Mathematics 10 (2022) 18, 3220, Universe 9 (2023) 431, Mon.Not.Roy.Astron.Soc. 538 (2025) 4, 3038-3041, arXiv:2503.10552*

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Outline

The Hubble tension

- General definitions & measurements

The Rotating Universe

- How this can be handled – a historical (re)view

The Model

- Spherical and cylindrical rotating dark matter

Results & Discussion

- Comparison to the standard cosmology & other models

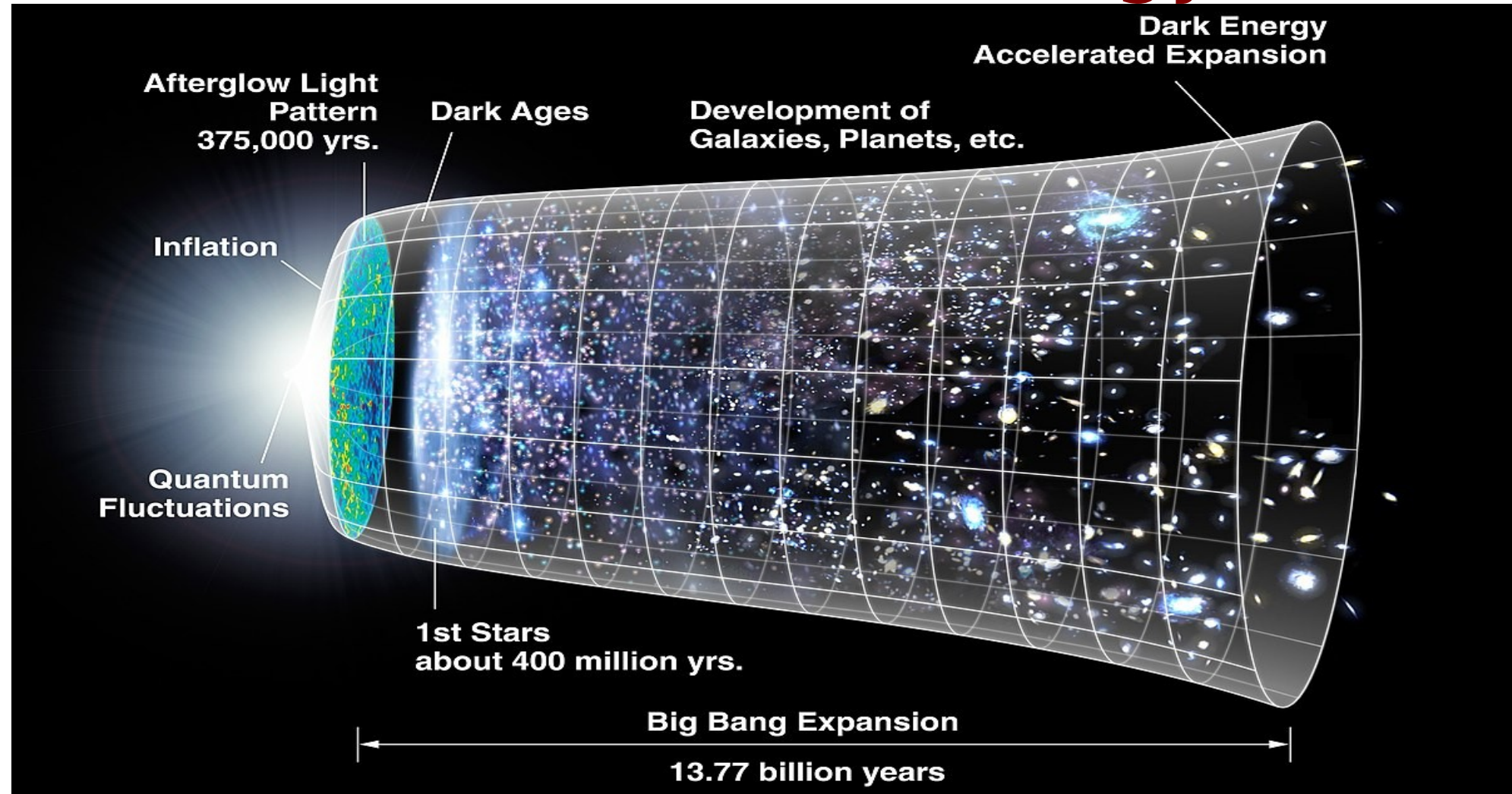
Conclusions

→ Yes, and πάντα κυκλoutai!



The Hubble Tension

Standard cosmology



The Hubble constant(?)

- **Measurement of Hubble-Lemaître:**

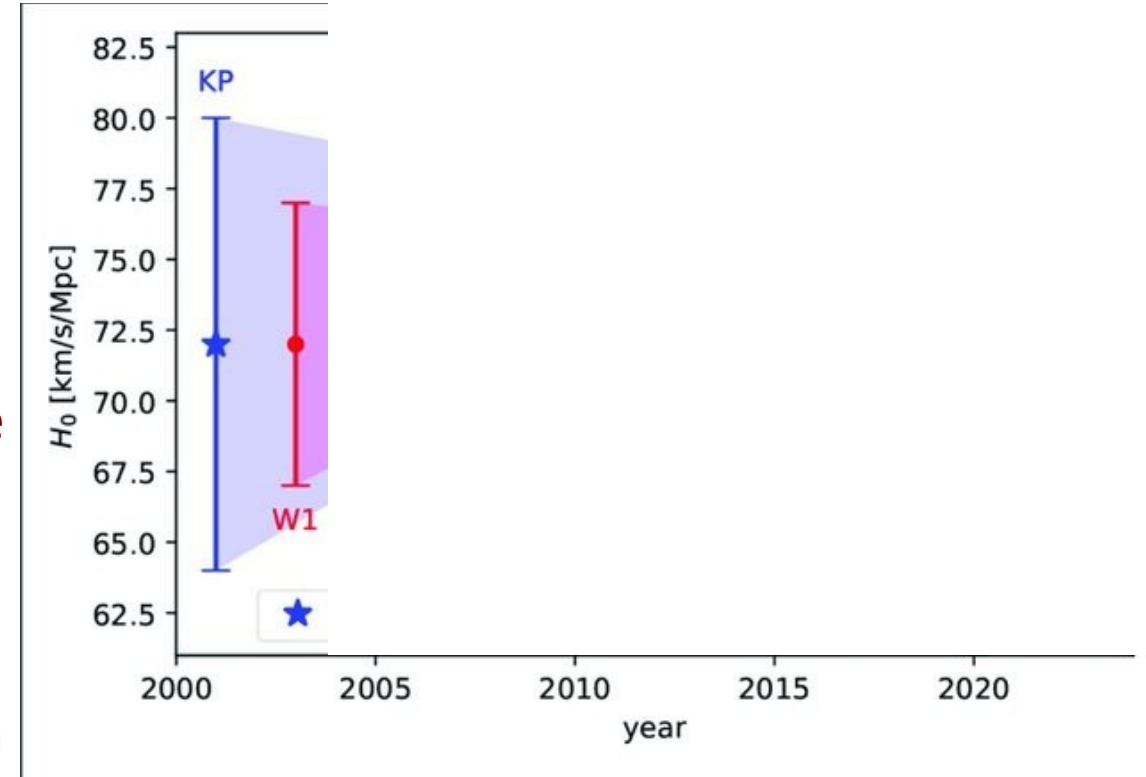
$$H^2 \equiv \left(\frac{\dot{a}}{a} \right)^2 = \frac{8\pi G}{3} \rho - \frac{kc^2}{a^2} + \frac{\Lambda c^2}{3}$$

- **Nearby Object's (Supernovae)**

This method relies on observations of Type Ia supernovae and other distance indicators.

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The CMB is the afterglow of the Big Bang, which can predict the expansion rate of the early universe and we can extrapolate it to the present day.



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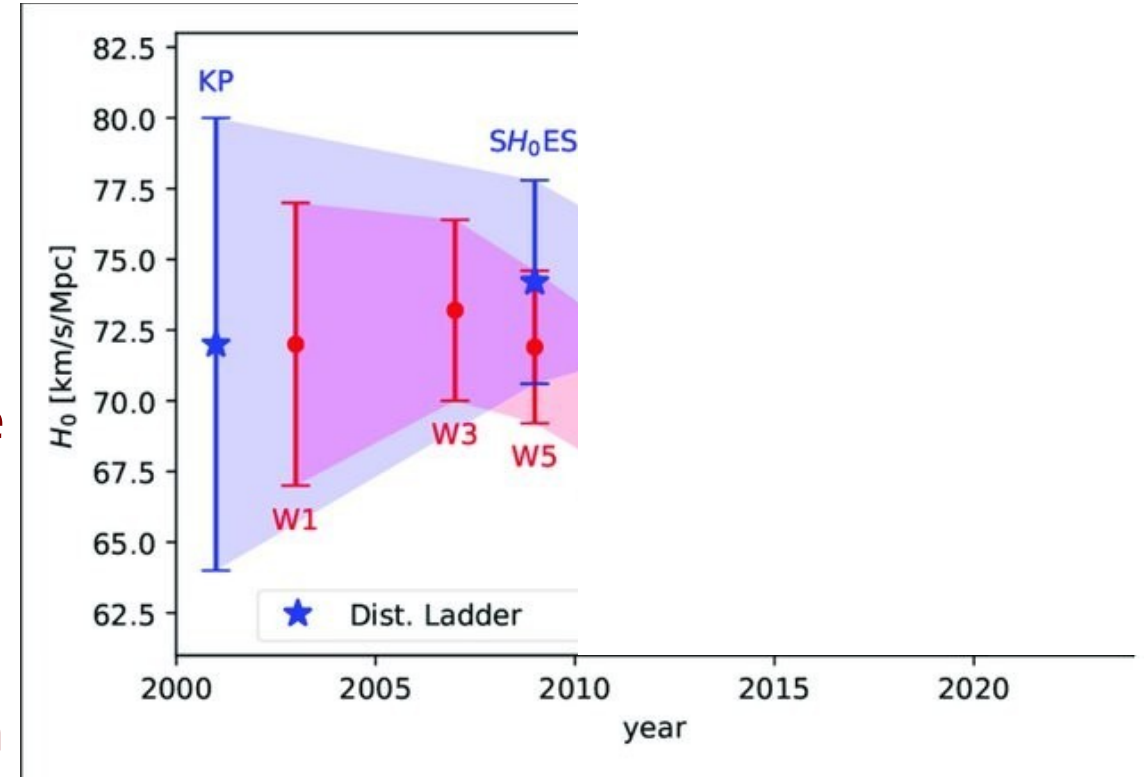
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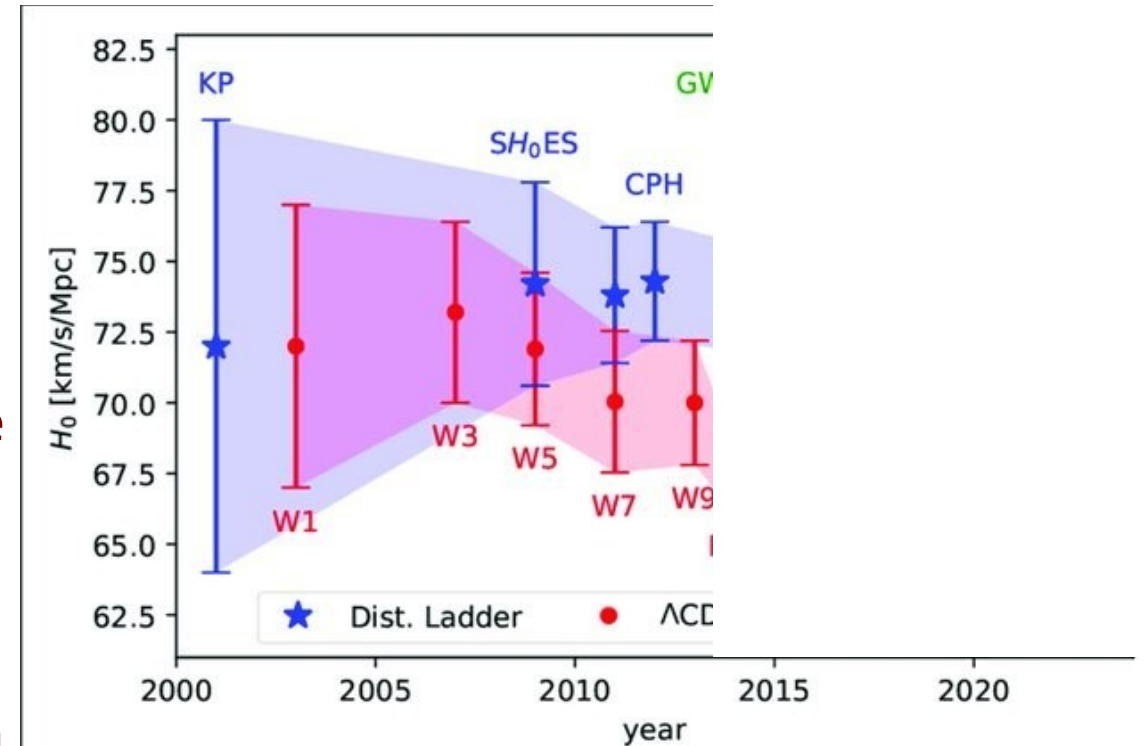
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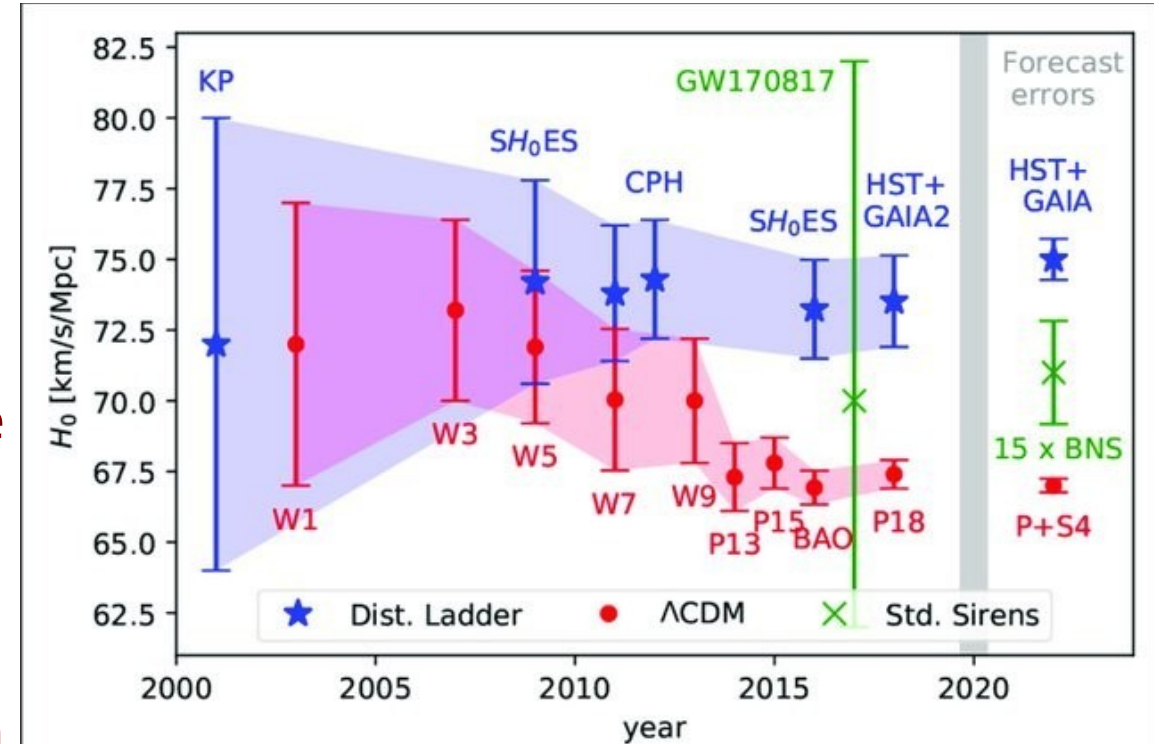
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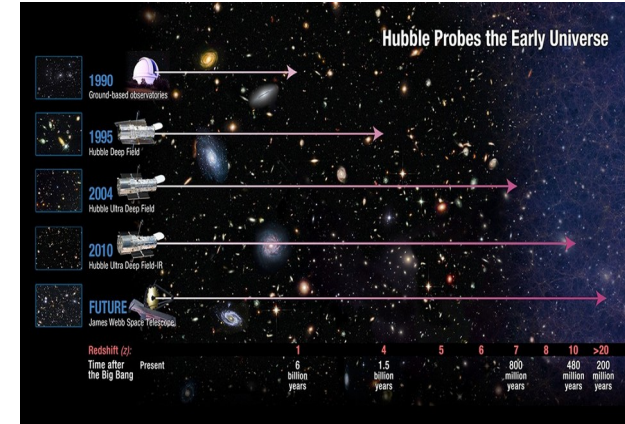
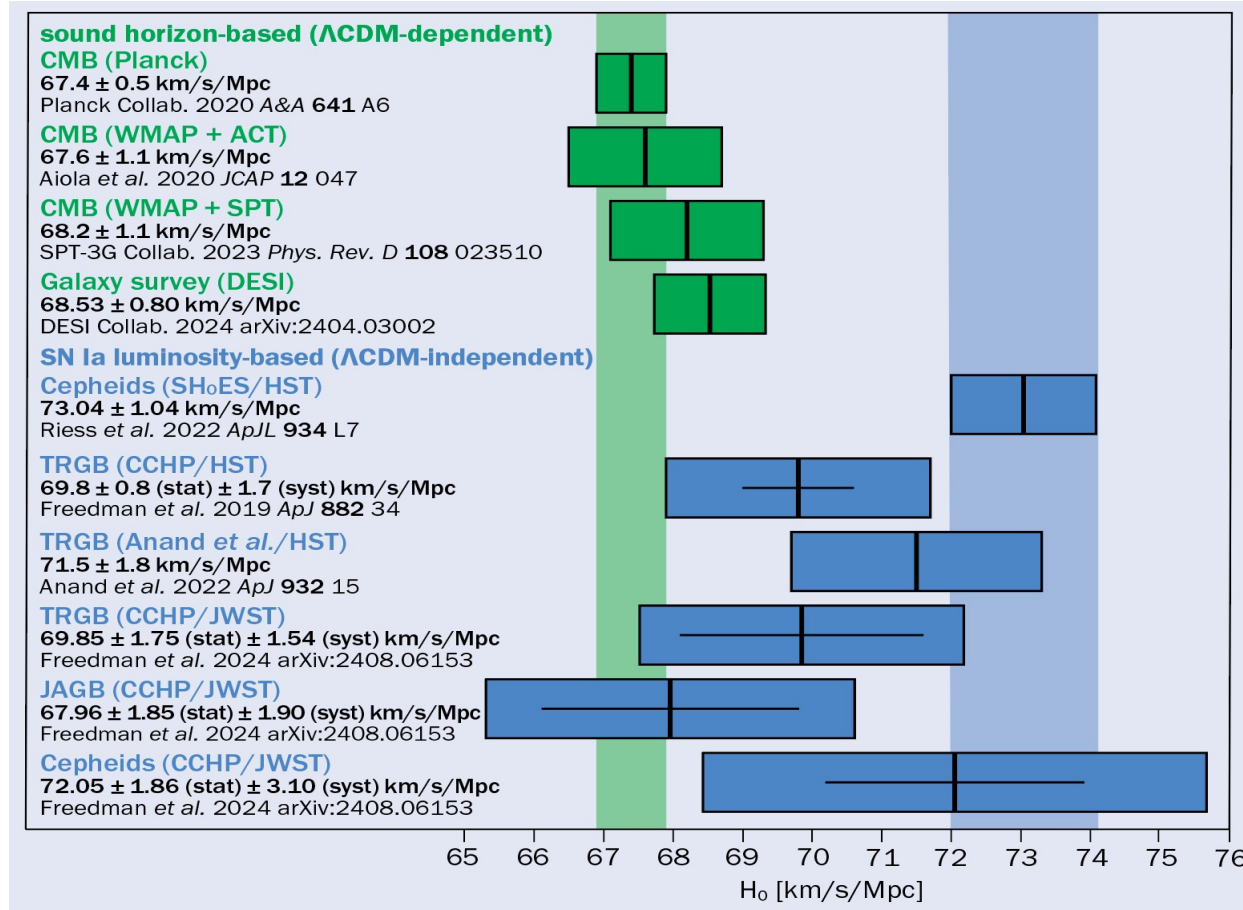
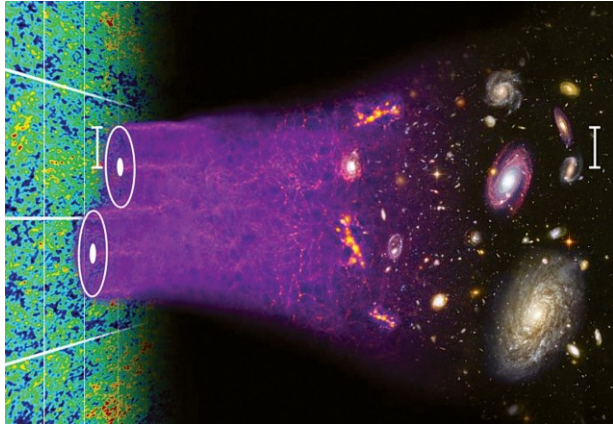
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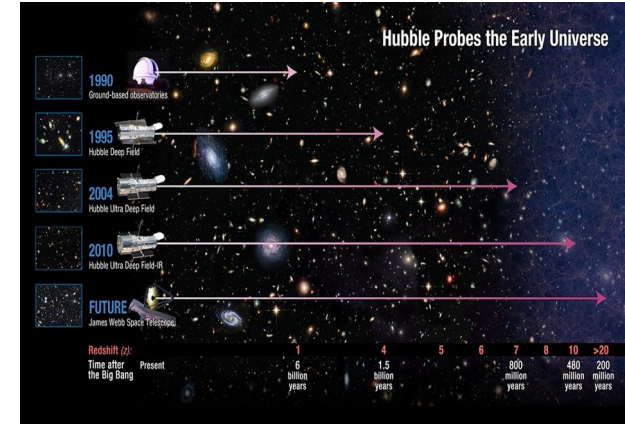
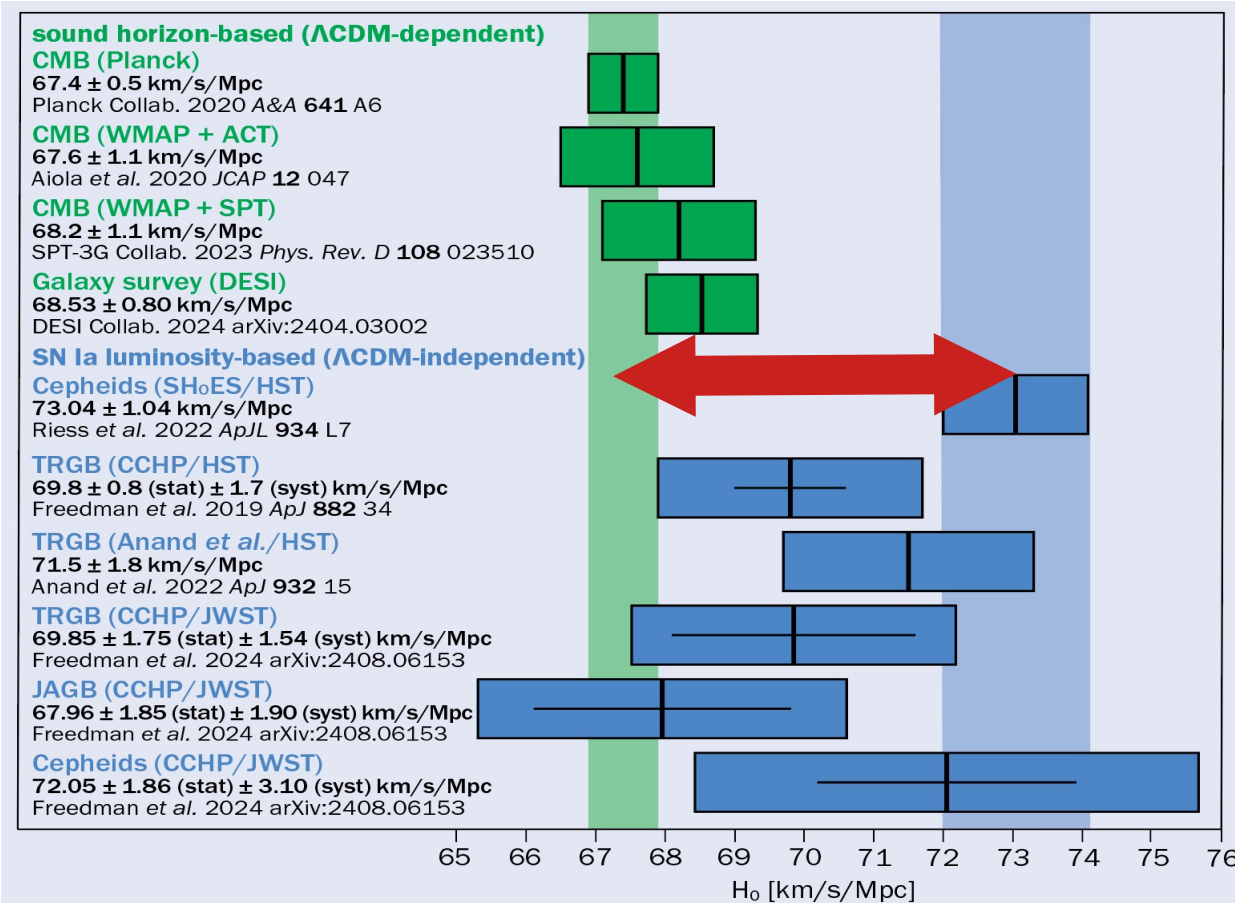
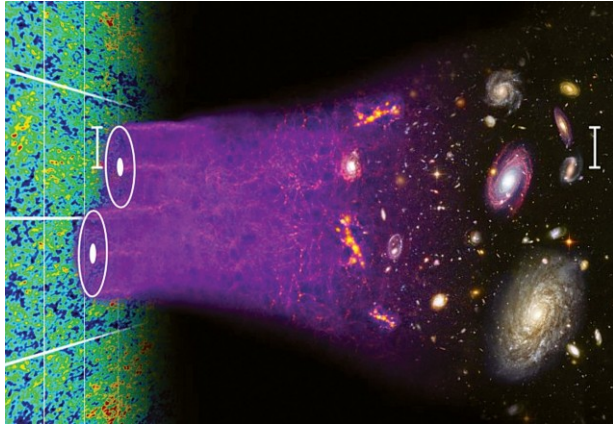
The Hubble tension



$$H_{\text{CMB}} = 67.4 \pm 0.5 \text{ kms}^{-1} \text{Mpc}^{-1}$$

$$H_{\text{SNe}} = 73.04 \pm 1.07 \text{ kms}^{-1} \text{Mpc}^{-1}$$

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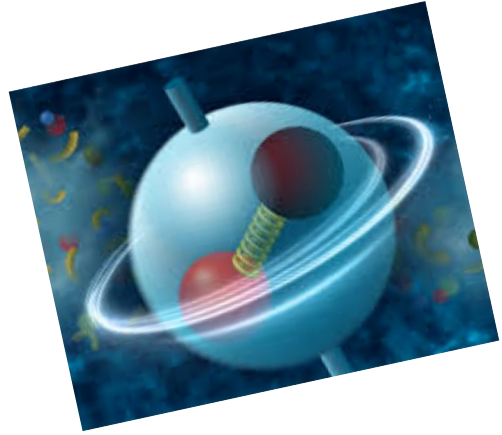
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How to resolve Hubble tension?

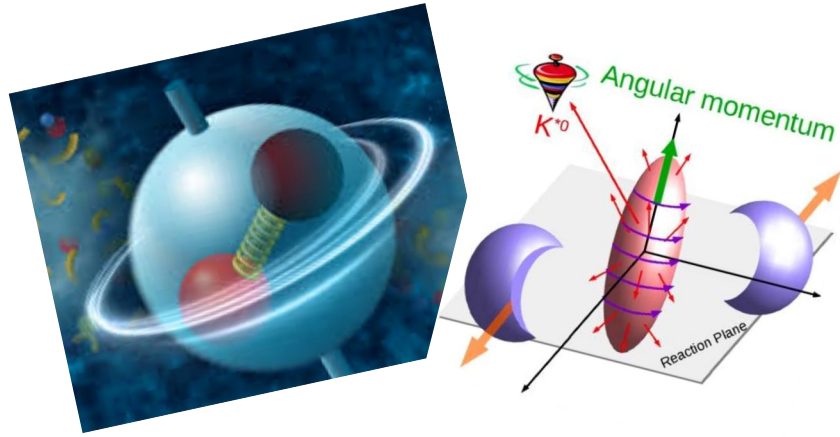
- **The Hubble tension suggests a potential problem with our understanding of the universe. Some possible explanations include:**
 - **Problems with our cosmological model:** The standard model of cosmology, Λ CDM, may be incomplete or incorrect.
 - **New Physics:** There could be unknown particles or forces influencing the universe's expansion.
 - **Systematic errors in measurements:** It's also possible that there are undetected errors in the measurement techniques.
- The Hubble tension is an active area of research, and scientists are working to refine measurements and explore potential solutions → **rotation of the Universe!**

The rotating Universe

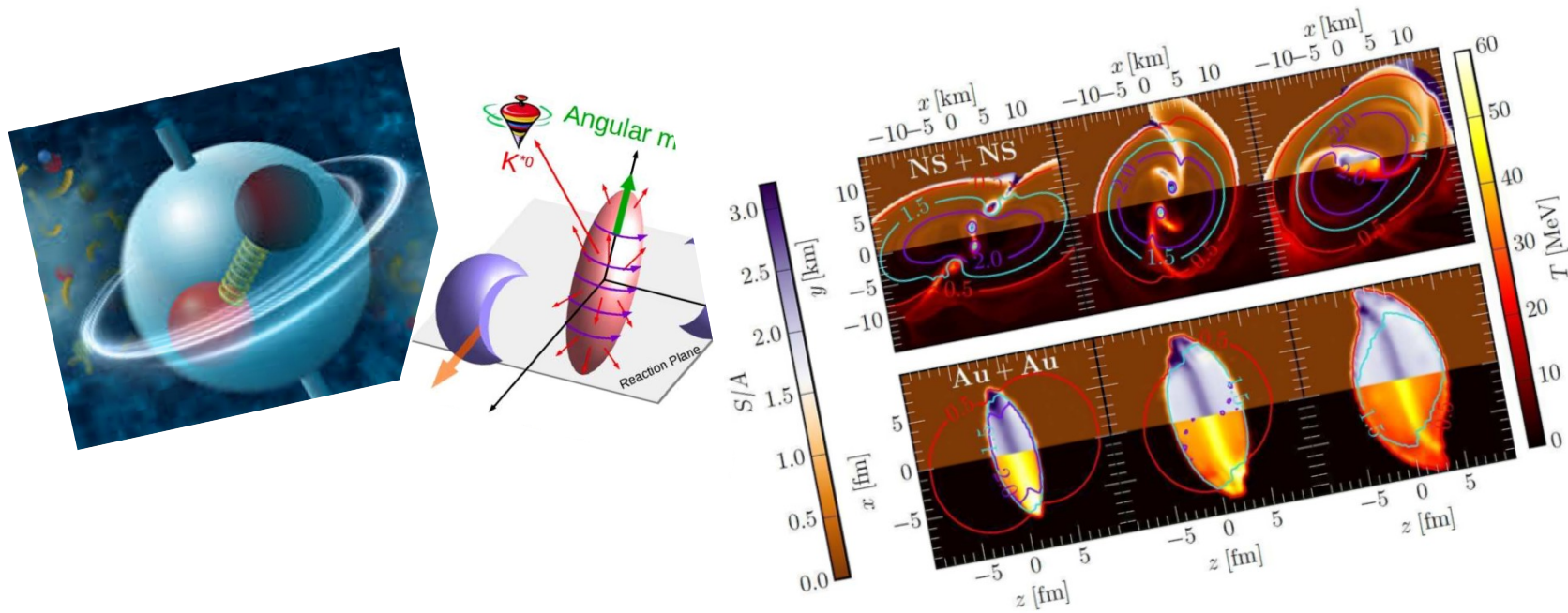
Are these rotating?



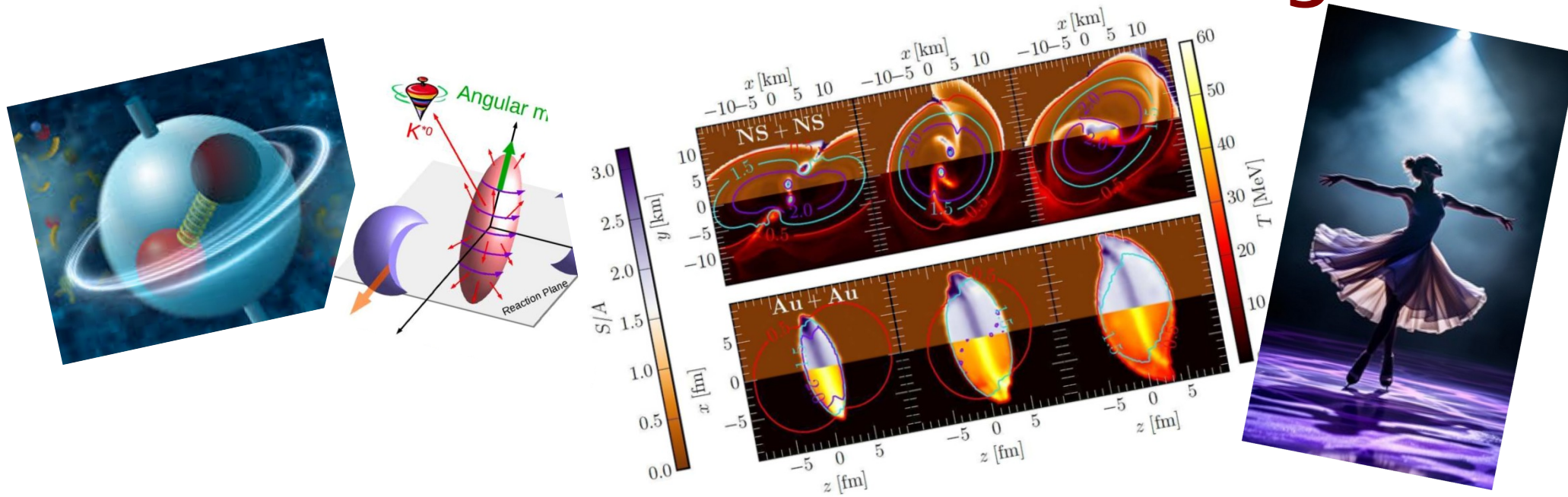
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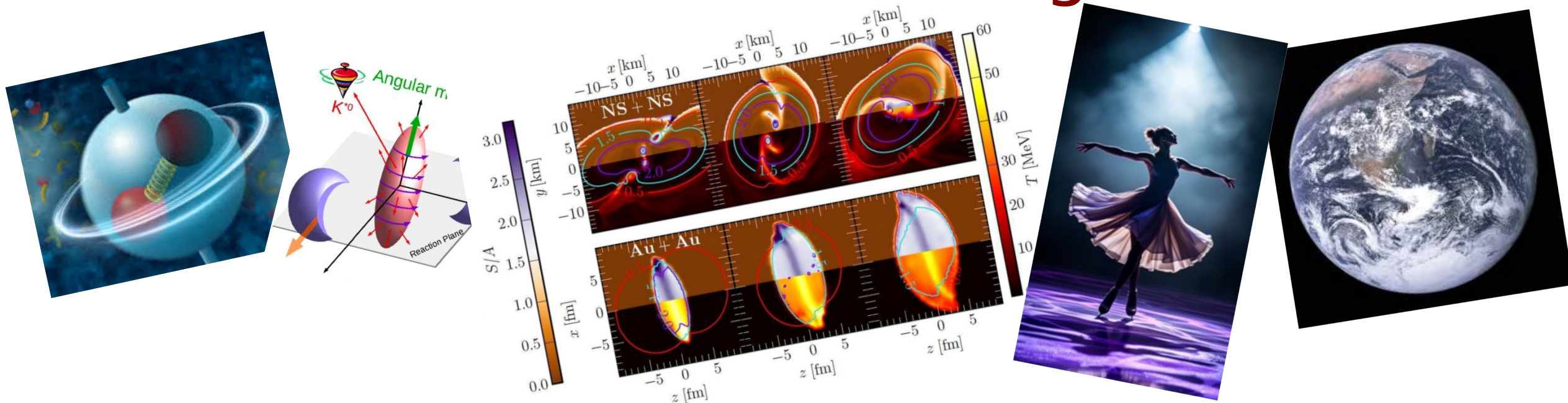
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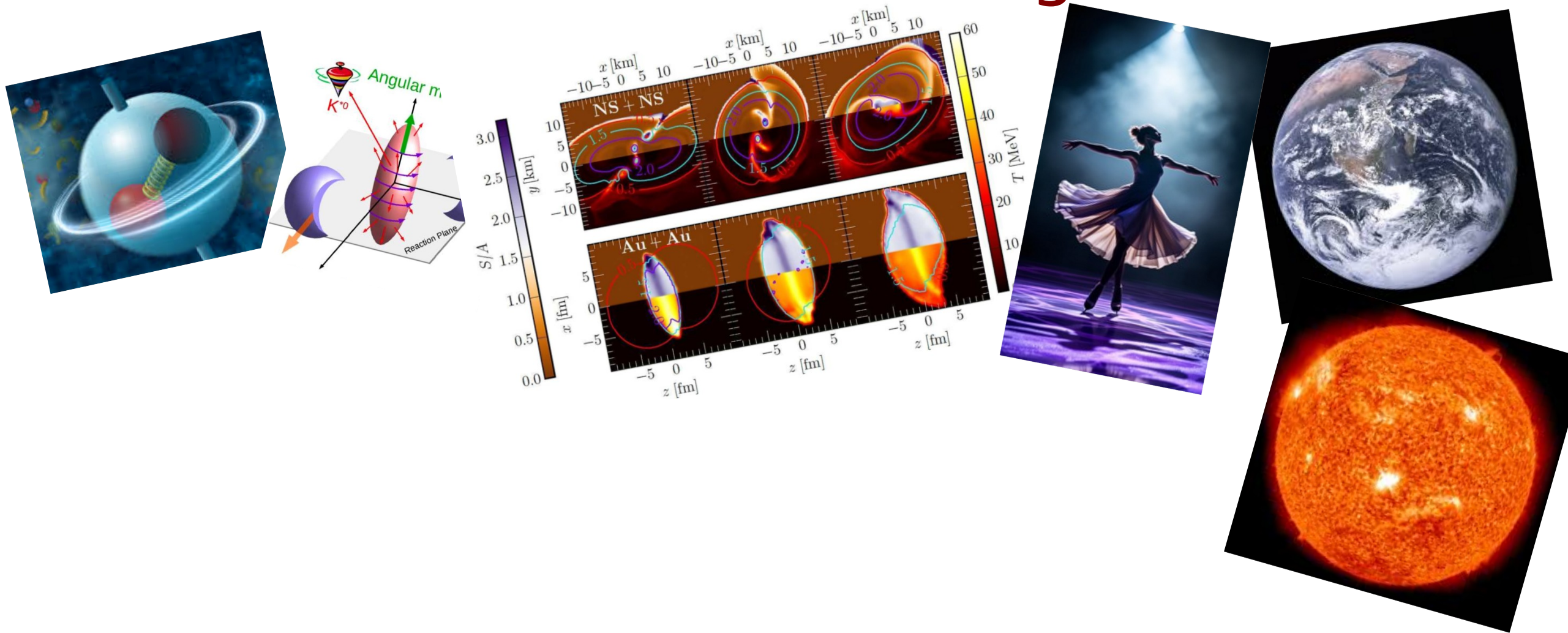
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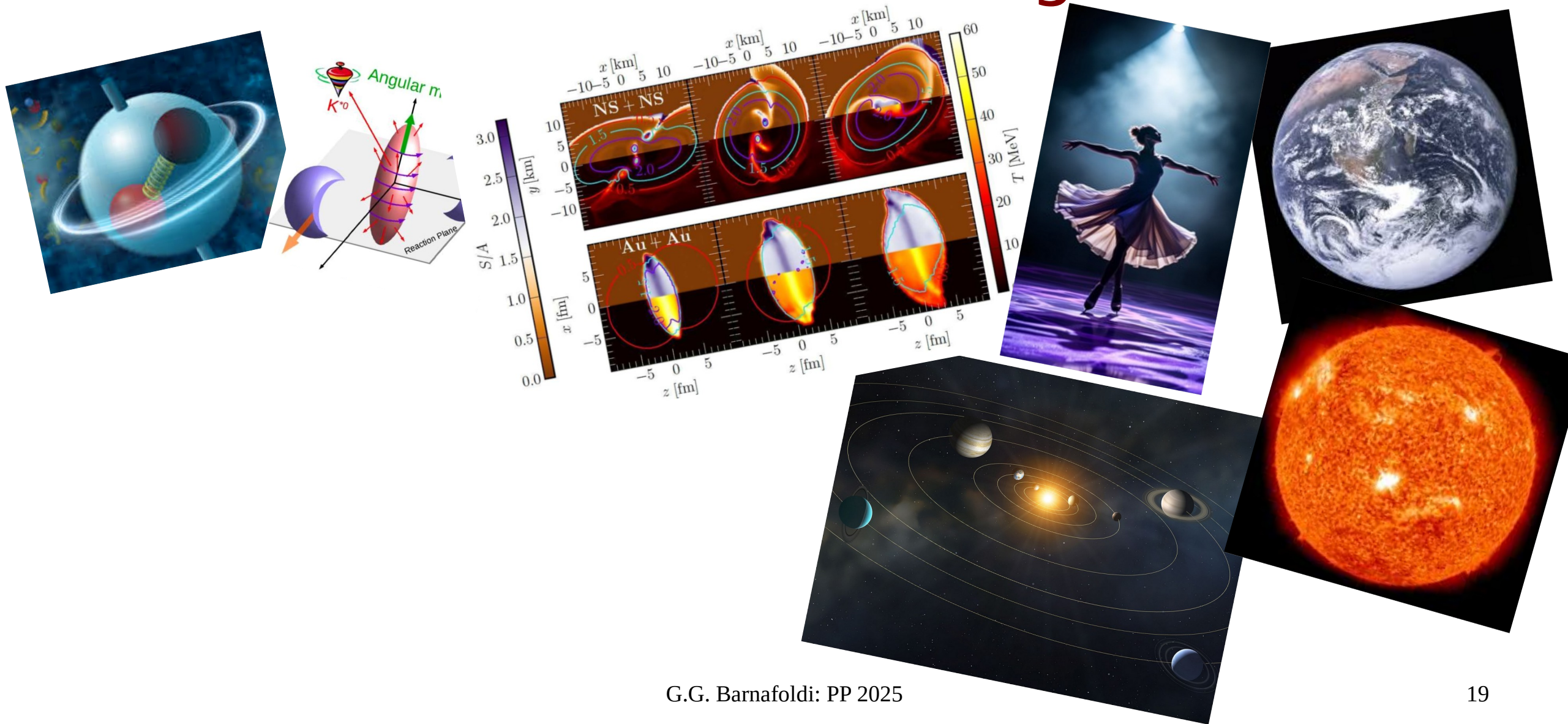
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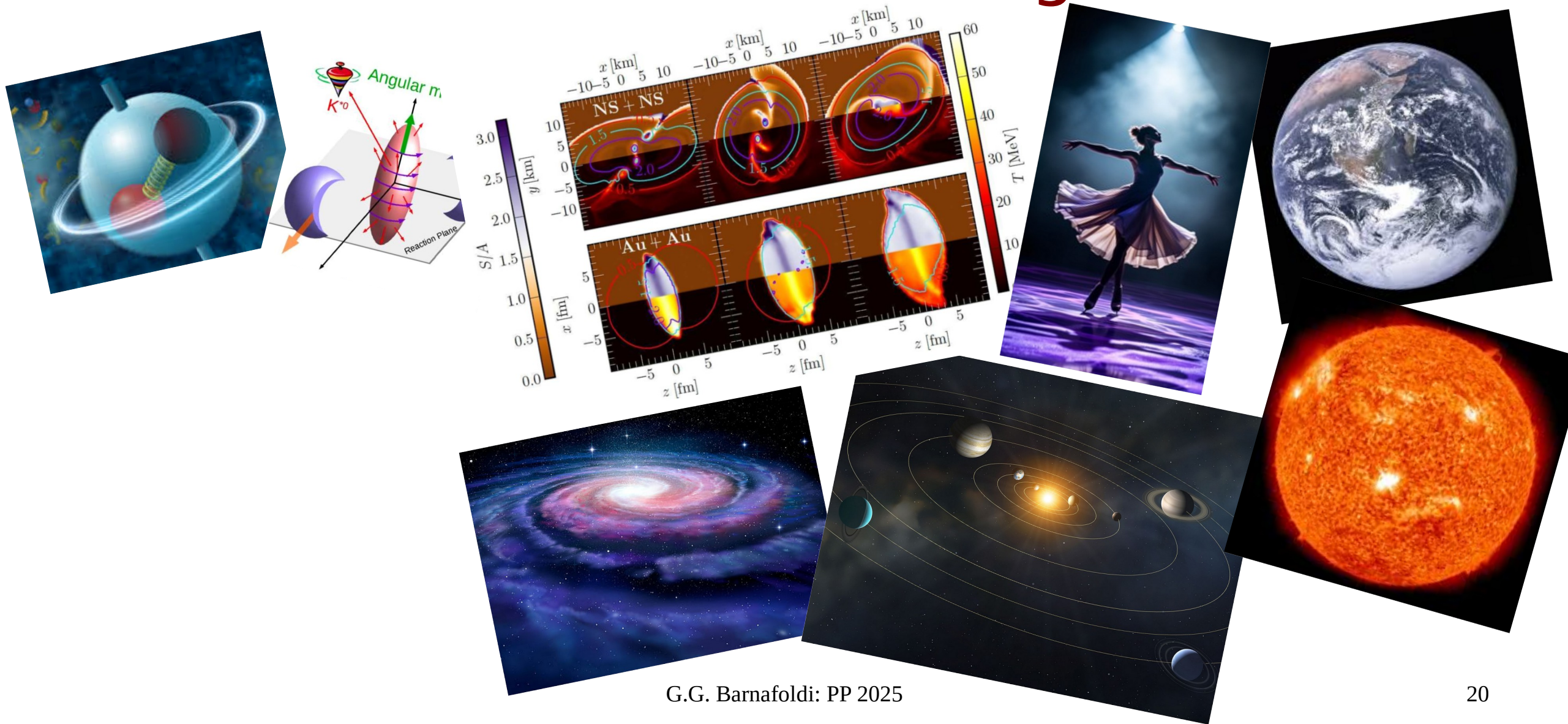
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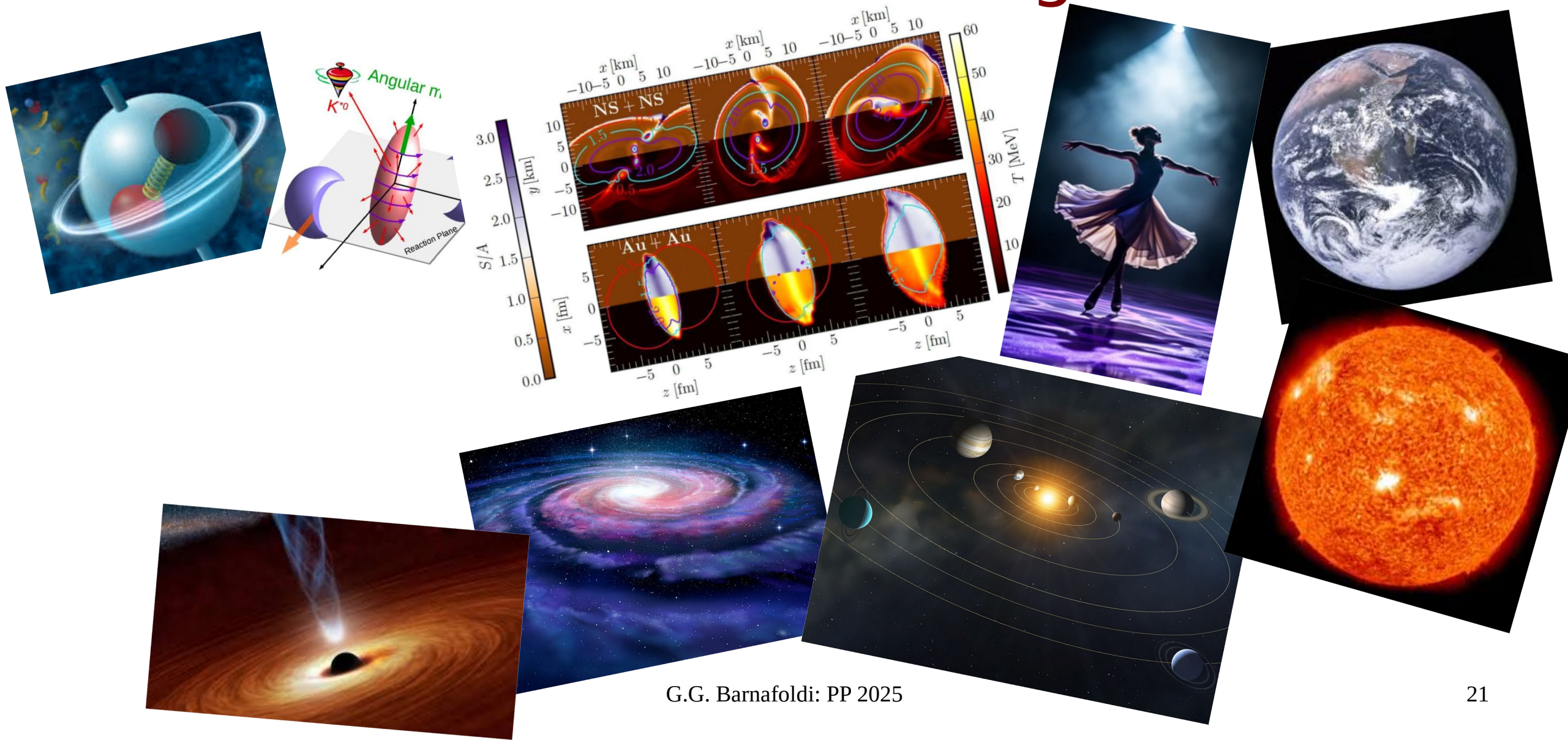
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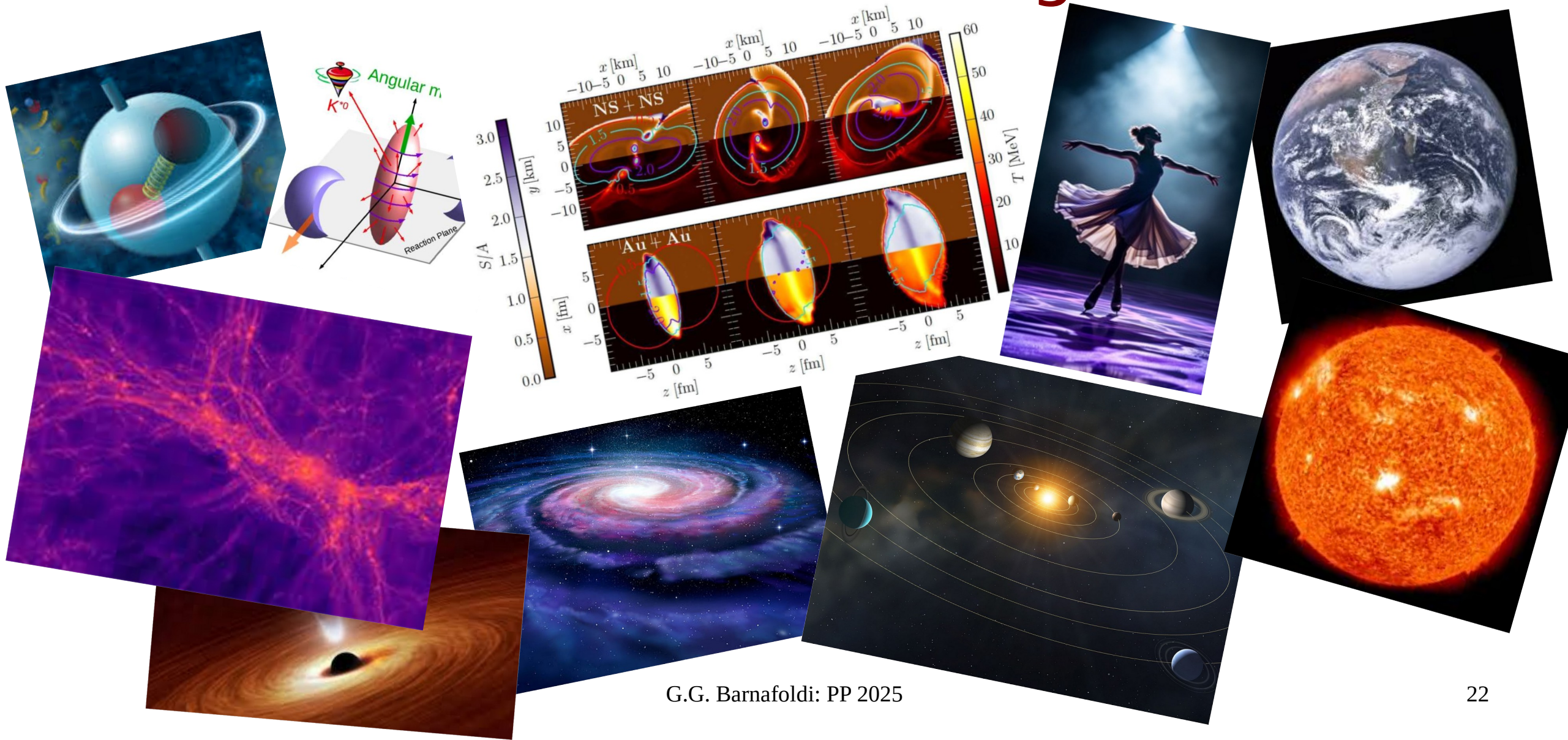
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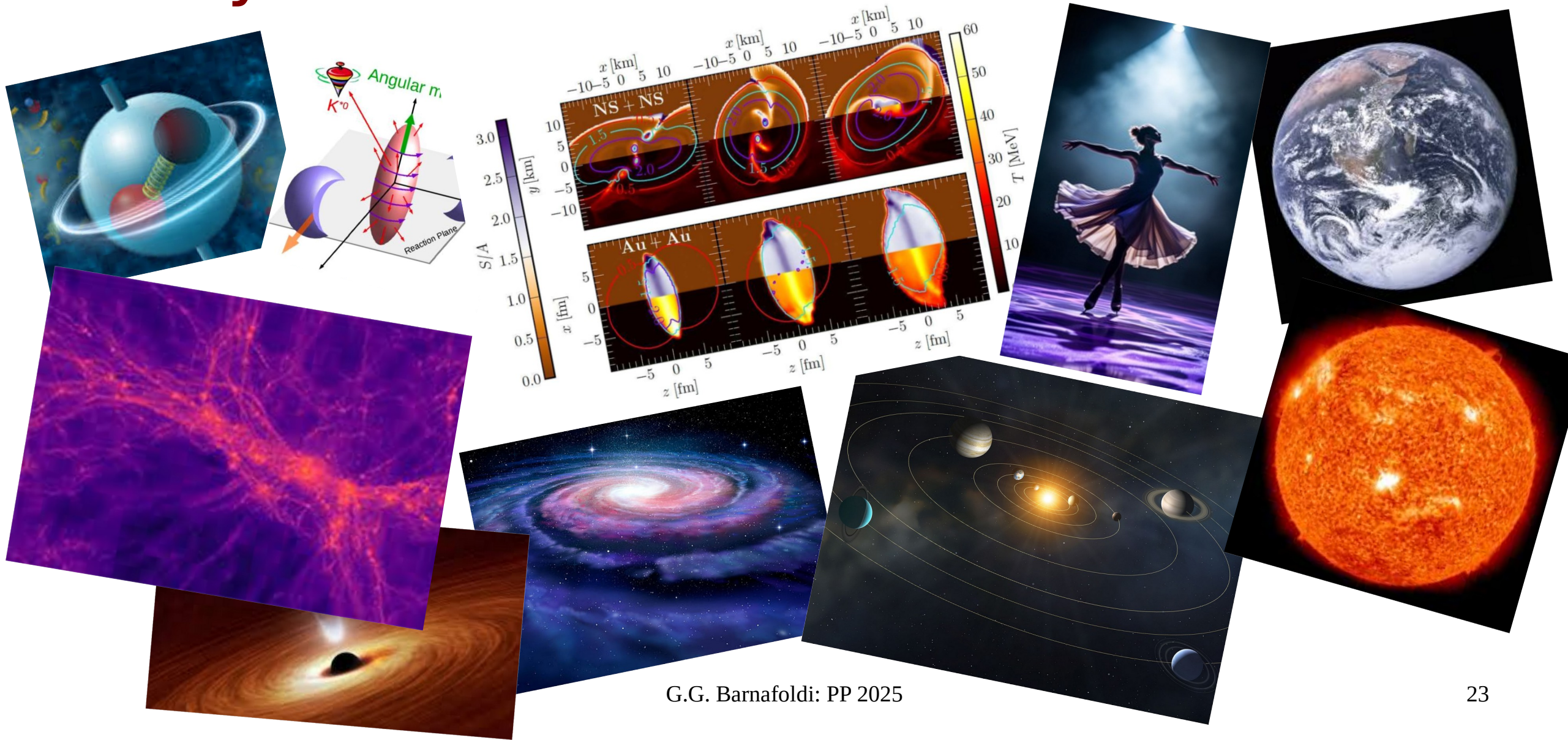
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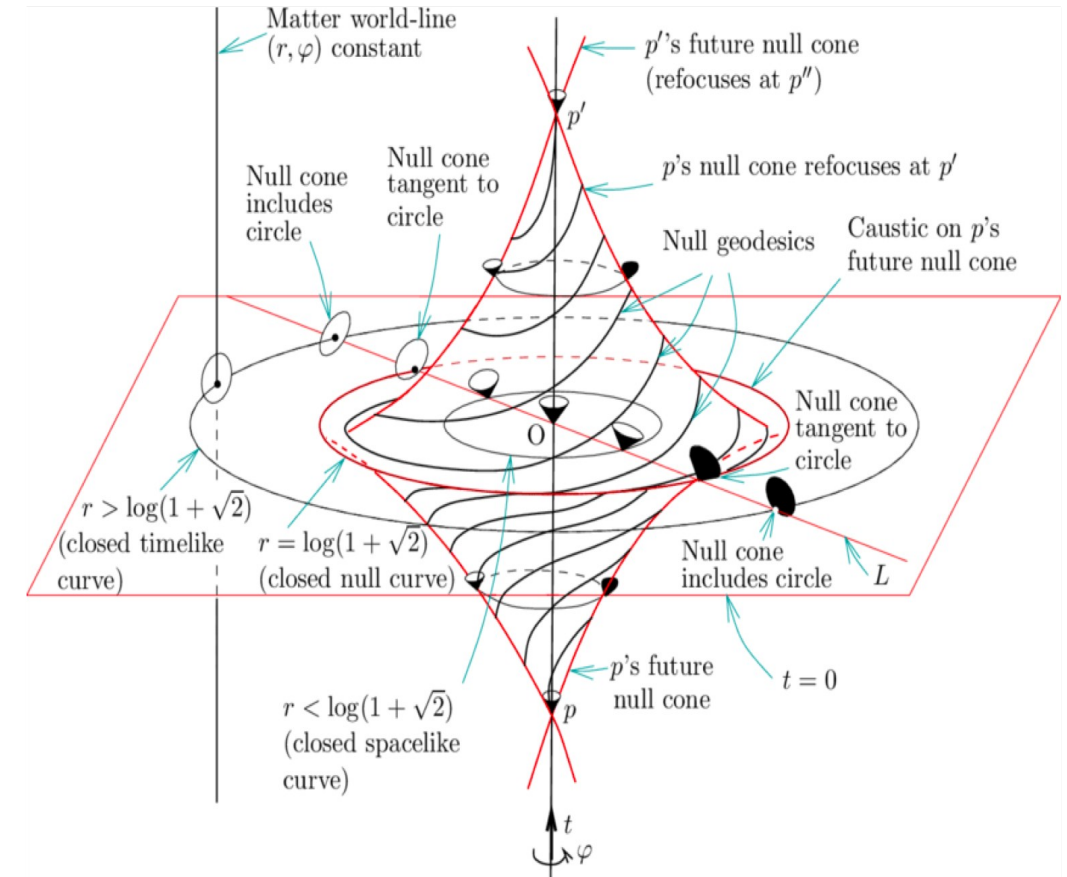
Why should not the Universe rotate?



Rotating cosmological models

- **Gödel inspired models** (1947): introduced a rotating universe followed by Heckmann & Schücking (1955-1961), later Silk (1966) and Hawking (1969). Obukhov (2000) made a generalized Gödel solution with global rotation. Buser, Kajari & Schleich (2013) visualized Gödel's universe). Anisotropies in a variety of Bianchi models with large vector perturbations corresponding to rotation are tightly constrained from Planck CMB data by Saadeh et al. (2016).
- **Other motivations for rotation:** Black Holes, spherically symmetric objects with horizons, display near maximal rotation as presented by Daly (2019). Anisotropic Hubble expansion in X-ray observations by Migkas et al. (2021) were suggested. A plausible syllogism is that the Universe has near-maximal rotation, motivated by cosmologies where the universe is the interior of a black hole (Pathria 1972).
- **There are many other proposed solutions** to the Hubble Puzzle (e.g. Di Valentino et al. 2021b) and any modification of the standard model expansion and (e.g. Knox & Millea: 2020). An average rotation effect has similar functional form as dark photons (Fabbrichesi, Gabrielli & Lanfranchi 2021; Aboubrahim et al. 2022), and (Cyr-Racine, Ge & Knox 2022).

→ Rotation affect the Hubble constant strongly....



The Model

The Model

- **Euler - Poisson equations without rotation:**

$$\partial_t \rho + \operatorname{div}(\rho \mathbf{u}) = 0 ,$$

$$\partial_t(\rho \mathbf{u}) + \operatorname{div}(\rho \mathbf{u} \otimes \mathbf{u}) = -\nabla P(\rho) - \rho \nabla \Phi + \rho \mathbf{g}$$

$$\nabla^2 \Phi = 4\pi G \rho ,$$

- **Dynamical parameters:** $\rho = \rho(r, t)$, $u = u(r, t)$, and $P = P(r, t)$

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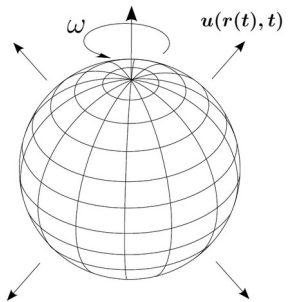
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- Spherical:**

$$\partial_t \rho + (\partial_r \rho)u + (\partial_r u)\rho + \frac{u\rho}{r} = 0 ,$$

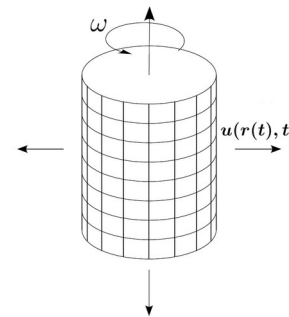


$$\partial_t u + (u \partial_r)u = -\frac{1}{\rho} \partial_r P - \partial_r \Phi(r) + g^*$$

$$\frac{1}{r} \frac{d}{dr} (r \partial_r \Phi) = 4\pi \rho .$$

- Cylindrical:**

$$\partial_t \rho + (\partial_r \rho)u + (\partial_r u)\rho + \frac{2u\rho}{r} = 0 ,$$



$$\partial_t u + (u \partial_r)u = -\frac{1}{\rho} \partial_r P - \partial_r \Phi(r) + g^*$$

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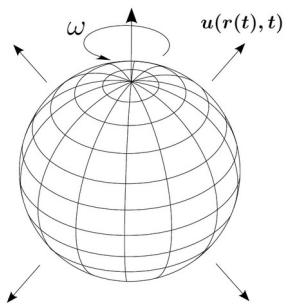
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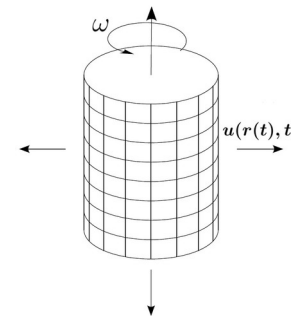


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$$\boxed{\frac{1}{r}} \frac{d}{dr} \left(r^{\boxed{2}} \partial_r \Phi \right) = 4\pi \rho ,$$

The EoS

- **A “stiff” dark matter EoS**

$$p = w\rho^n \quad \text{for} \quad n = 1 \quad \longrightarrow \quad \frac{dp}{d\rho} = c_s^2 = w$$

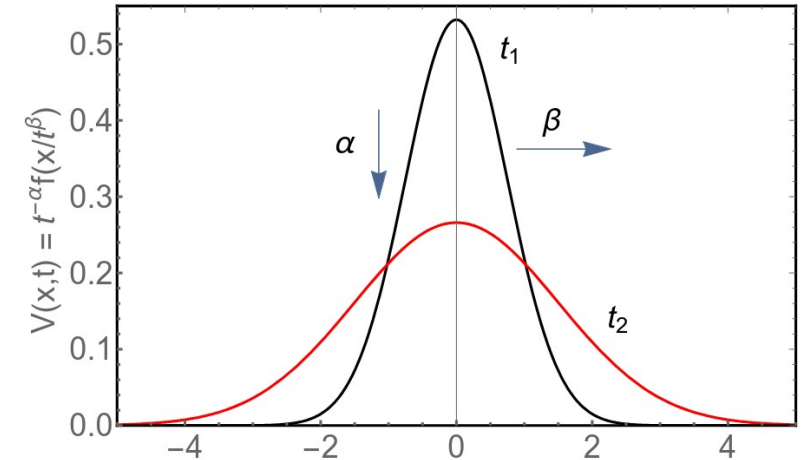
- **Speed of sound values - in general [Mathematics 10 (2022) 18 3220]**

- (i) $w = 0$ means the EoS for ordinary non-relativistic ‘matter’ (e.g., cold dust);
- (ii) $w = 1/3$ means ultra-relativistic ‘radiation’ (including neutrinos) and, in the very early universe, other particles that later become non-relativistic;
- (iii) $w = -1$ is the simplest case and describes the expanding universe, hypothetical phantom energy $w < -1$ would cause Big Rip;
- (iv) $w \neq -1$ means quintessence as hypothetical fluid;
- (v) $w = -1/3$ is responsible for the flatness of the Big Bang;
- (vi) A scalar field ϕ can be viewed as a sort of perfect fluid with EoS of $w = \frac{\frac{1}{2}\phi_t^2 - U(\phi)}{\frac{1}{2}\phi_t^2 + U(\phi)},$

Self-similar solutions

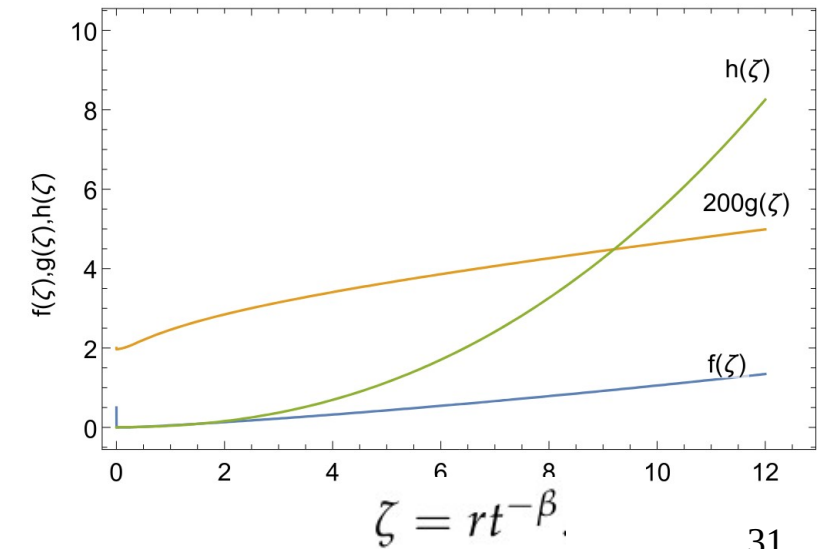
- **A Sedov-von Neumann-Taylor *ansatz*:**

- $V(x, t) = t^{-\alpha} f\left(\frac{x}{t^\beta}\right) := t^{-\alpha} f(\omega)$.
- Similarity parameters
- Gaussian (regular) heat conduction: $\alpha = \beta = 1/2$;



- **Shape functions (f, g, h) :**

$$\left. \begin{aligned} u(r, t) &= t^{-\alpha} f\left(\frac{r}{t^\beta}\right) \\ \rho(r, t) &= t^{-\gamma} g\left(\frac{r}{t^\beta}\right) \\ \Phi(r, t) &= t^{-\delta} h\left(\frac{r}{t^\beta}\right) \end{aligned} \right\} \zeta = rt^{-\beta}.$$



Self-similar solutions

- **Applying the Sedov-von Neumann-Taylor *ansatz*:**

- Partial Differential Equation (PDE) system becomes Ordinary Differential Equation System (ODE) for the shape functions, of the $\zeta = rt^{-\beta}$.

$$-\zeta g'(\zeta) + f'(\zeta)g(\zeta) + f(\zeta)g'(\zeta) + \frac{2f(\zeta)g(\zeta)}{\zeta} = 0,$$

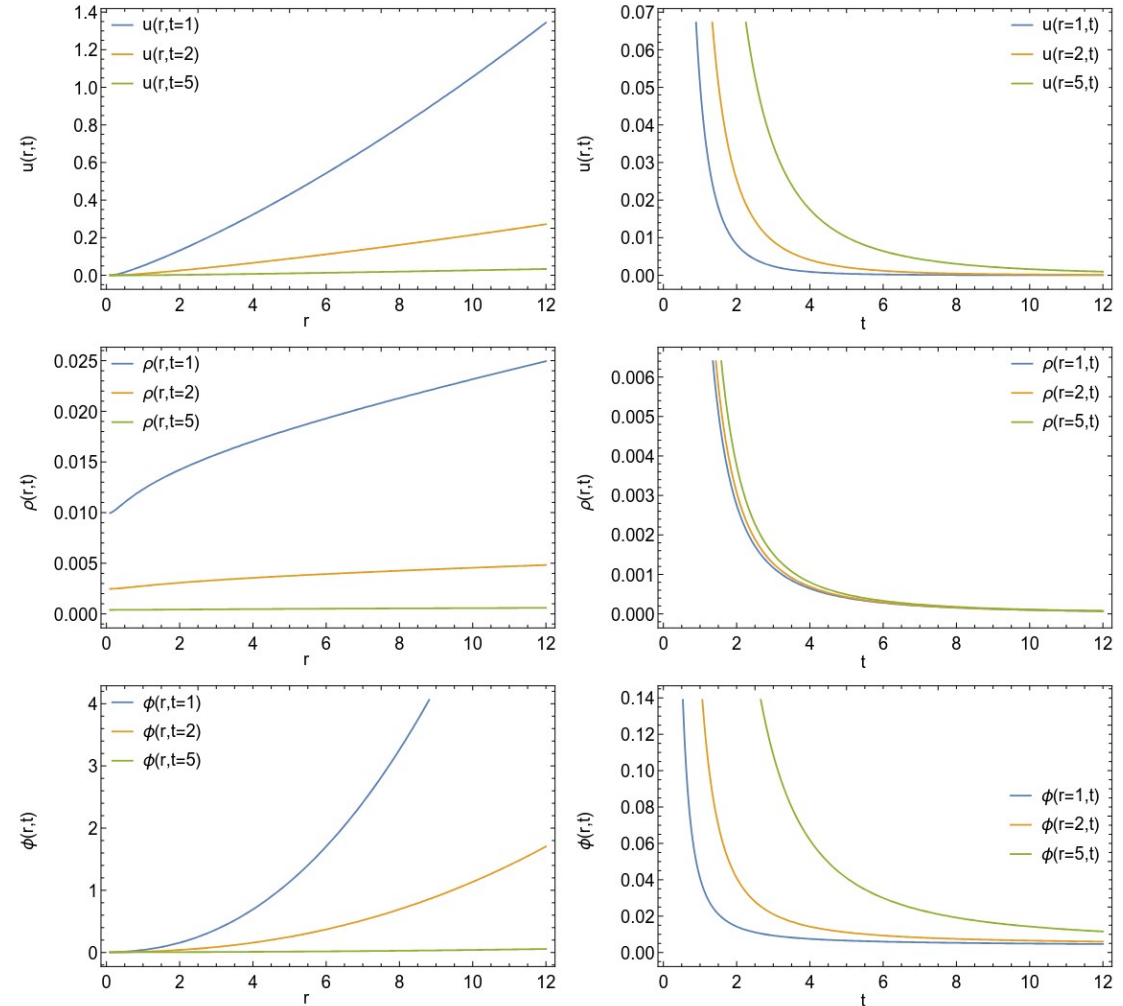
$$-\zeta^2 f'(\zeta) + \zeta f'(\zeta)f(\zeta) = -\frac{wg'(\zeta)}{g(\zeta)} - h'(\zeta)\zeta + \omega^2 \sin \theta \zeta^2,$$

$$h'(\zeta) + h''(\zeta)\zeta = g(\zeta)4\pi G\zeta.$$

- **Similarity exponents for these are:** $\alpha = 0, \beta = 1, \gamma = 2$, and $\delta = 0$.

Self-similar solutions

- **Non-rotating case:**
 - Radial velocity and density profile decreases with time
 - Density and velocity profiles are different in radius, but both are increasing.
 - Gravitational potential are decreasing hyperbolically with time, and becomes asymptotically flat.
- **See in [Universe 9 (2023) 431]**

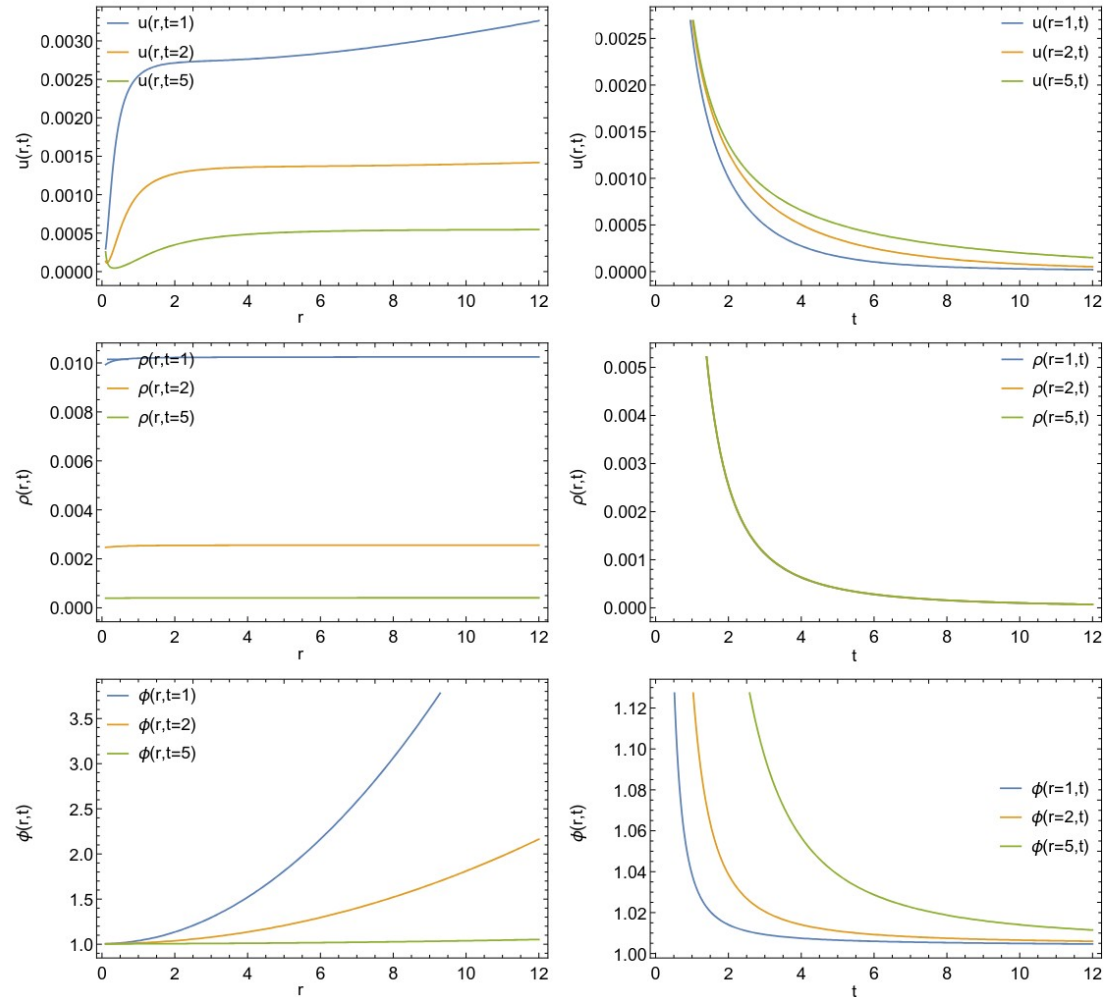


Self-similar solutions

- **Rotating case:**

- Radial velocity and density profile decreases with time
- Density and velocity profiles are different in radius: velocity curves have plateau and density become flat.
- Gravitational potential are decreasing hyperbolically with time, and becomes asymptotically flat.

- **See in [Universe 9 (2023) 431]**

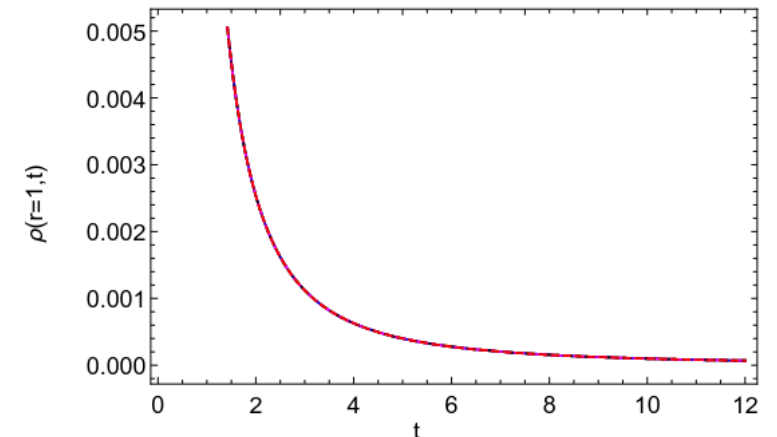
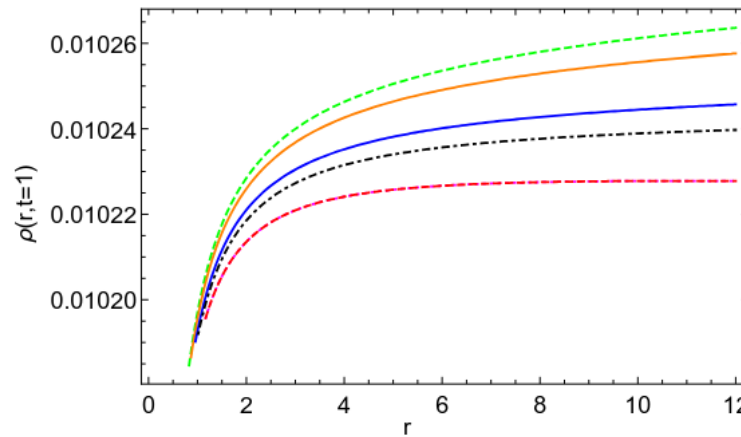
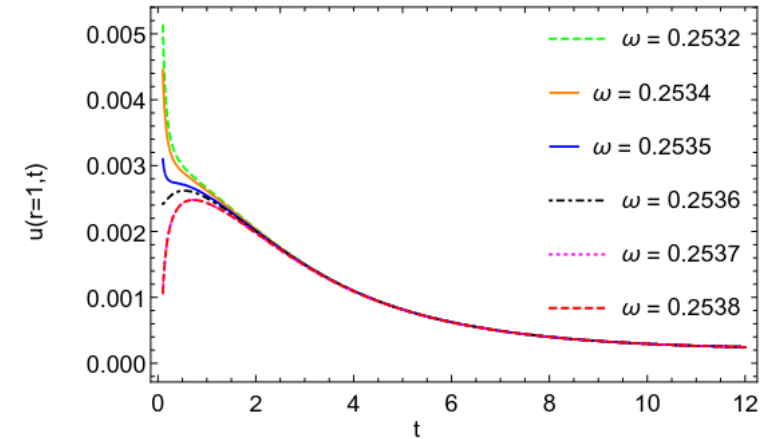
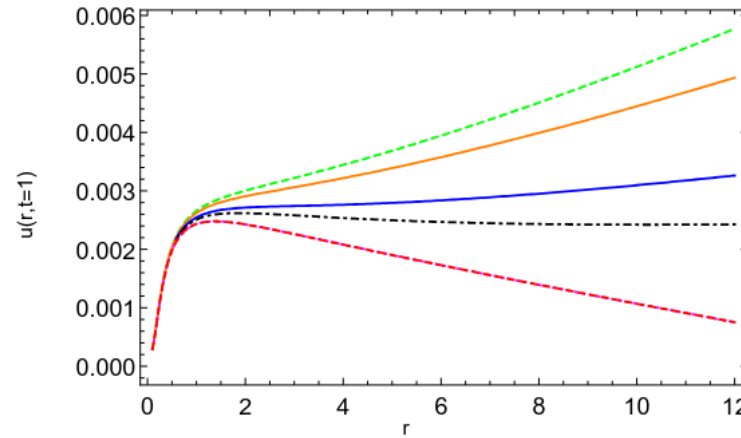


Self-similar solutions

- **Rotating case:**

- Radial velocity and density as a function of time at different rotation have nice asymptotic limit.
- Both velocity and density profiles are changing well with rotation.
- Maximal rotation exists certainly.

- **See in** [Universe 9 (2023) 431]



Resolving the Hubble tension

Connection to Hubble parameter

- **Expansion rate & the Hubble parameter:**

- Relative distance with the scale: $R(t) = a(t) l$

- Assuming the mass conservation:

$$\frac{d}{dt}M(t) = 4\pi \frac{d}{dt} \int \rho(a(t)l, t) a^3(t) l^2 dl \stackrel{!}{=} 0 \longrightarrow \rho(a(t)l, t) \frac{d}{dt} \left[\rho(a(t)l, t) \right] = -3 \frac{\dot{a}(t)}{a(t)}$$

- Kinematical condition:

$$\frac{d}{dt}R(t) = u(R(t), t) \Rightarrow \frac{1}{g(R(t), t)} \frac{d}{dt} \left[t^{-\gamma} g(R(t), t) \right] = -3 \frac{t^{-\alpha} f(R(t), t)}{R(t)}$$

with the solution:

$$R(t) = u_1 t^{\beta + \frac{\gamma}{\kappa}} \exp \left[-\frac{3u_1 t^\mu}{\mu\kappa} \right] \times \left[3^{-\frac{\gamma}{\mu\kappa}} u_2 t^{\gamma/\kappa} \left(\frac{u_1 t^\mu}{\nu} \right)^{-\frac{\gamma}{\mu\kappa}} \Gamma \left(1 + \frac{\gamma}{\nu}, \frac{3u_1 t^\mu}{\nu} \right) - \mathcal{H}_1 u_1 \right]^{-1}.$$

Connection to Hubble parameter

- Expansion rate & the Hubble parameter:**

- In the solution we assume a Taylor exp:

$$u(r, t) \sim u_1 \zeta^1 + u_2 \zeta^2$$

then, relative distance can be given:

$$R(t) = \frac{t}{\mathcal{H}_1 t^{\frac{3u_1-2}{\kappa}} + \frac{3u_2}{2-3u_1}}, \quad \text{where } \kappa = \frac{6}{7}$$

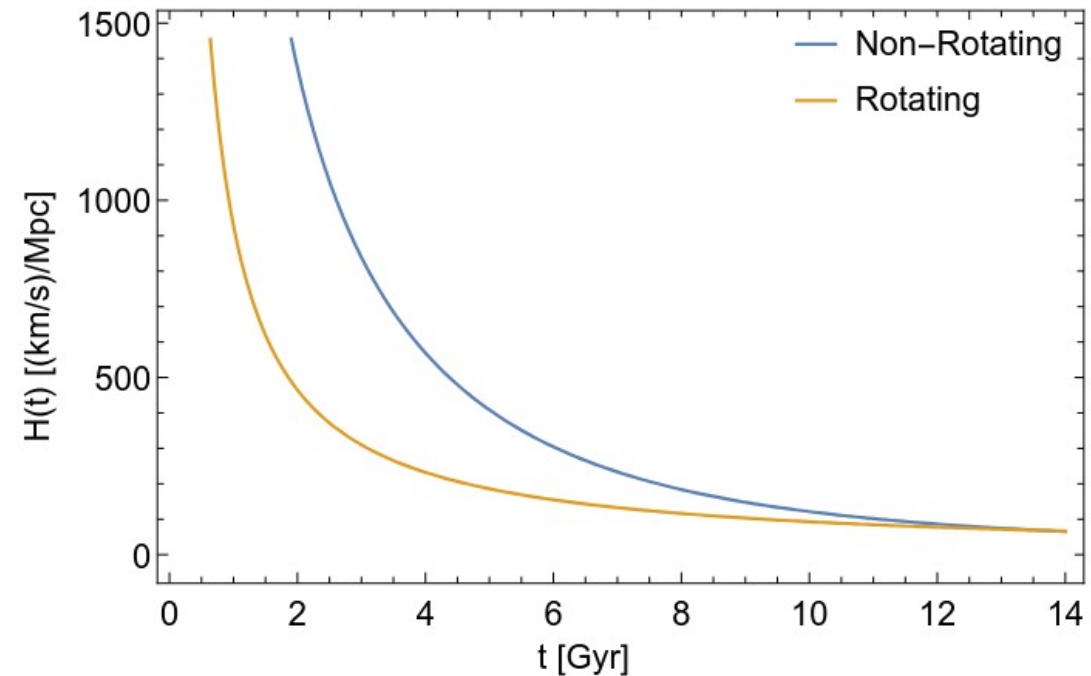
- From the integration constant:

$$\left. \frac{\dot{a}(t)}{a(t)} \right|_{t=t_0} = H_0, \text{ if } a(t_0) = 1 \quad \longrightarrow$$

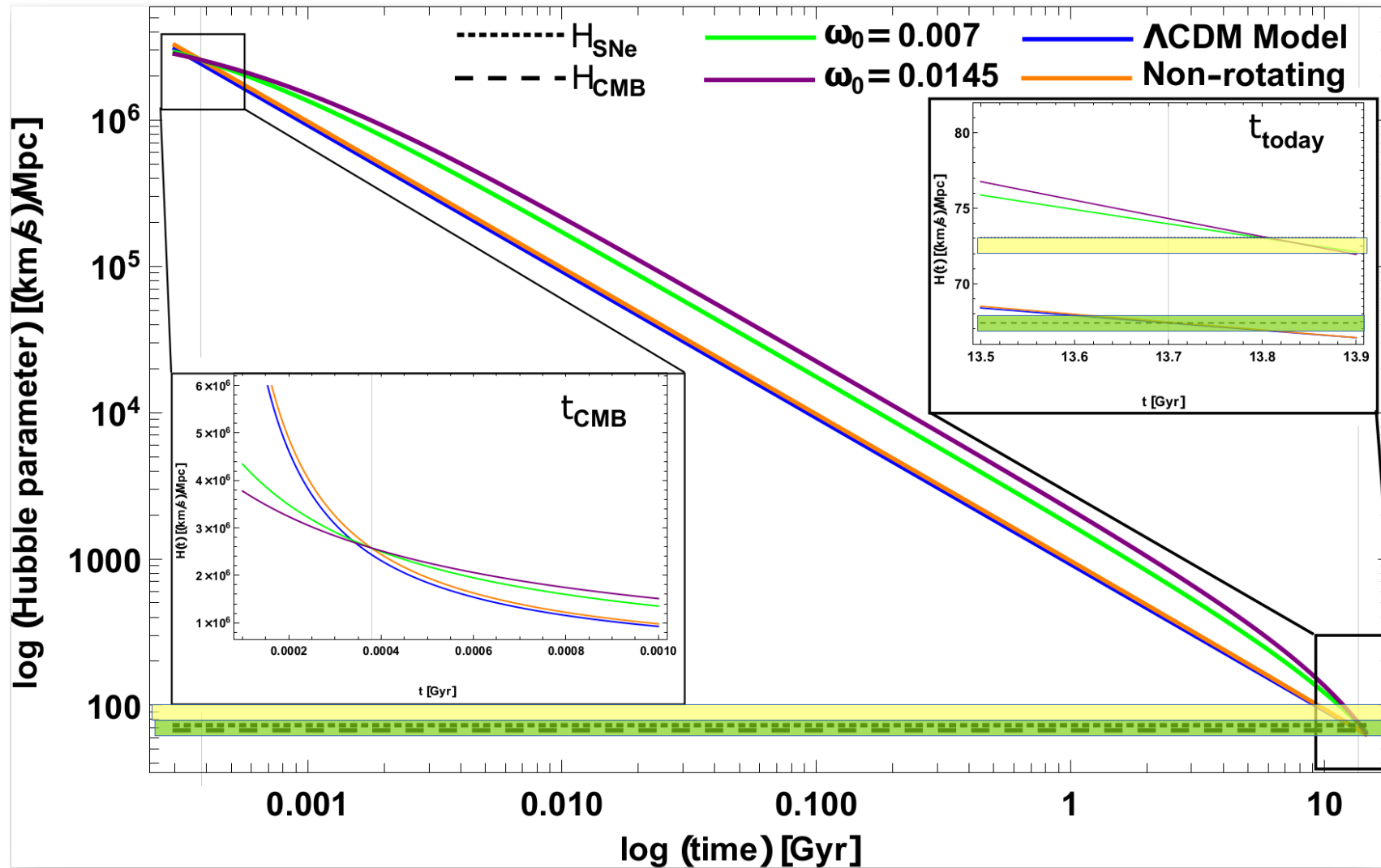
- For Euclidian (flat) Universe (today):

$$H_0 = 66.6^{+4.1}_{-3.3} \text{ km/s/Mpc}$$

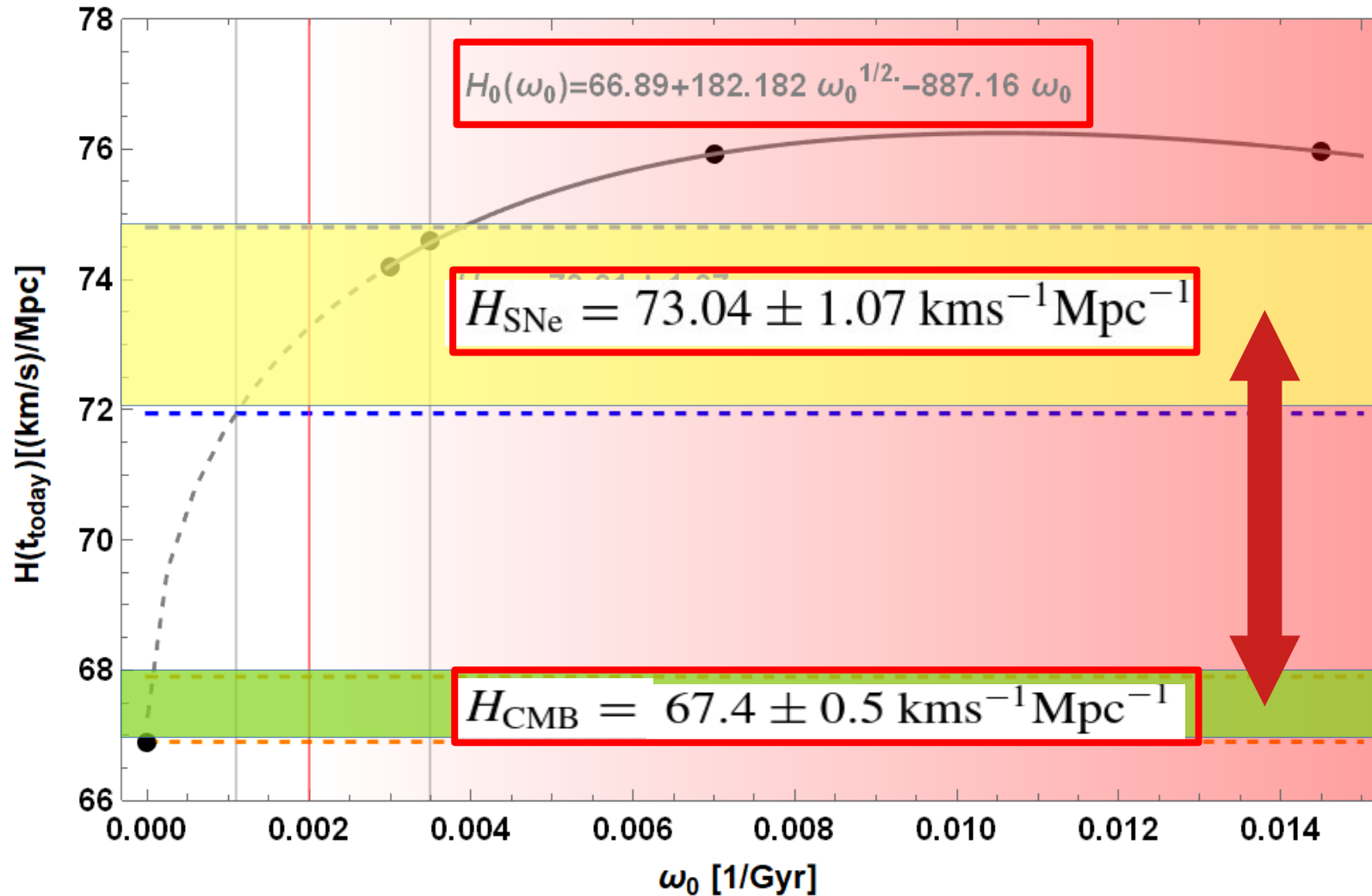
$$\left. \frac{\Omega_M(t)}{\Omega(t)} \right|_{t=t_0} \sim \left. \frac{E_{kin}(R(t), t)}{E_{tot}(R(t), t)} \right|_{t=t_0} = 0.26$$



The calculated Hubble parameter



Hubble parameter in rotating Universe



Discussion of the rotation

- **The rotation of the Universe:**

- Today, the rotation of the Universe is: $\omega_0 = 0.002^{+0.001}_{-0.0009} \text{ Gyr}^{-1}$
- Rotation at the CMB, this value was: $\omega(t_{\text{CMB}}) = 3.54^{+1.3}_{-1.2} \text{ Myr}^{-1}$
- Since angular rotation approximately $|\omega(t)| = \omega_0 a^{-2}(t)$ and assuming the speed should be below the speed of light $\omega \lesssim H$, then taking $H(a) \sim a^{-3/2}$, one can estimate the rotation limit today:
→ The result is the maximal rotation, which is below the observational limit, but also compatible with many rotational models:

$$\omega_0 \lesssim H_0 a^{1/2}(t_{\text{eq}}) \simeq 0.002 \text{ Gyr}^{-1}$$

Summary

- **παντα κυκλoutai!**

- Classical Newtonian model with dark matter with Sedov – von Neumann – Taylor self similarity
- The Universe is rotating at maximal angular velocity today: 500 billion year/rotation, which resolves the Hubble tension
- Compatibility with measurements and other theoretical approaches

- **News & youtube videos:**

- Video1, Video2, Anton Petrov Video, Rakéta, NewsWeek, ScienceAlert, Iflscience, ScienceAlert, Studyfinds, Rakéta - report, Cosmos Magazine, AstroBite, ScienMag, Message to Eagle, The Universe Today, Brian Koberlein, Enholm, EarthSky, Studyfinds, The Debrief, SpaceWeekly



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πάντα κυκλoutai!

πάντα ρεῖ!



G.G. Barnafoldi: PP 2025